

Impact of Velocity Space Distribution on Hybrid Kinetic-MHD Simulation of the (1, 1) Internal Kink Mode

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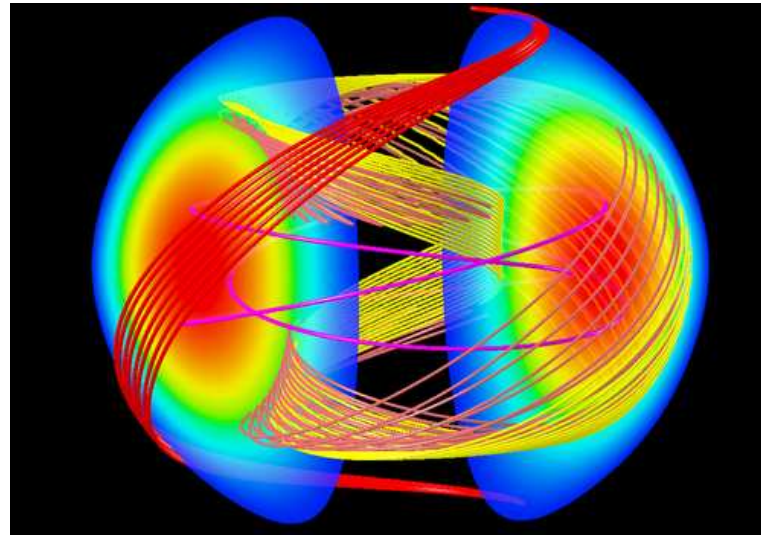
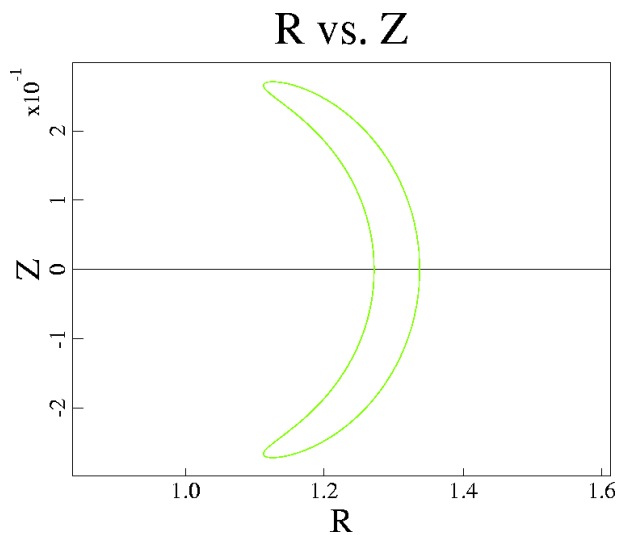
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U. Washington

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Hybrid Kinetic-MHD - Building Bridges

- capture kinetic effects lost in MHD approximation
- lost kinetic effects can **significantly** effect MHD instabilities^a
 - high energy tail of plasma executes non-trivial orbits - large gyro-orbits, banana-orbits, potatoes



- implement δf PIC in FE MHD code (NIMROD)

^aW. Park, et al, "Three-dimensional hybrid gyrokinetic- magnetohydrodynamic simulation", *Physics of Fluids B*, 4, 1992

Numeric representation of NIMROD fields

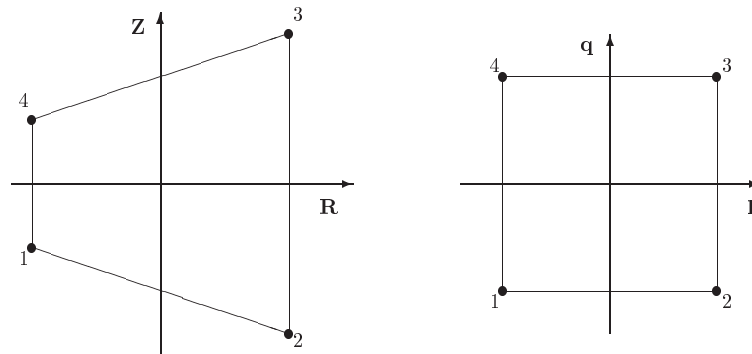
- the **perturbed** NIMROD fields are in FE-Fourier representation

$$\delta A(\mathbf{x}, t) = \sum_j A_{j,0}(t) \alpha_{j,0} + \sum_j \sum_n (A_{j,n}(t) \alpha_{j,n} + c.c.)$$

where

$$\alpha_{j,n} = N_j(p, q) \exp(in\phi)$$

(p, q) are logical coordinates



- structured finite elements
- PIC becomes nontrivial in FE

PIC in FEM

- gather/scatter require logical coordinates
- FE representation for (R, Z) need to be inverted

$$R = \sum_j R_j N_j(p, q), \quad Z = \sum_j Z_j N_j(p, q),$$

- Use Newton method to solve for (p, q) given (R, Z)

$$\begin{pmatrix} p^{k+1} \\ q^{k+1} \end{pmatrix} = \begin{pmatrix} p^k \\ q^k \end{pmatrix} + A(p_k, q_k) \begin{pmatrix} R - R^k \\ Z - Z^k \end{pmatrix}$$

- Jacobian gives gradient of shape functions for free
- sorting is also required for parallelization
- efficient load balancing is tricky
- current implementation is limited to linear elements for PIC
 - high order elements for fluid variables

The Hybrid Kinetic-MHD Equations^a

- in the limit $n_h \ll n_0$, $\beta_h \sim \beta_0$, quasi neutrality, only modification of MHD equations is addition of the **hot particle pressure tensor** in the momentum equation:

$$\rho \left(\frac{\partial \mathbf{U}}{\partial t} + \mathbf{U} \cdot \nabla \mathbf{U} \right) = \mathbf{J} \times \mathbf{B} - \nabla p_b - \nabla \cdot \underline{\mathbf{p}}_h$$

the subscripts b, h denote the bulk plasma and hot particles

- the steady state equation

$$\mathbf{J}_0 \times \mathbf{B}_0 = \nabla p_0 = \nabla p_{b0} + \nabla p_{h0}$$

- evolved momentum equation is ($\mathbf{U}_s = 0$)

$$\rho_s \frac{\partial \delta \mathbf{U}}{\partial t} = \mathbf{J}_s \times \delta \mathbf{B} + \delta \mathbf{J} \times \mathbf{B}_s - \nabla \cdot \delta \underline{\mathbf{p}}_b - \nabla \cdot \delta \underline{\mathbf{p}}_h$$

^aC.Z.Cheng, 'A Kinetic MHD Model for Low Frequency Phenomena', *J. Geophys. Rev* **96**, 1991

Deposition of $\delta \underline{\mathbf{p}}_h$ onto Finite Element grid

- assume CGL-like form $\delta \underline{\mathbf{p}}_h = \begin{pmatrix} p_{\perp} & 0 & 0 \\ 0 & p_{\perp} & 0 \\ 0 & 0 & p_{\parallel} \end{pmatrix}$
- evaluate pressure moment at a position \mathbf{x} is

$$\begin{aligned} \delta p(\mathbf{x}) &= \int m(v - V_h)^2 \delta f(\mathbf{x}, \mathbf{v}) d^3v \\ &= \sum_{i=1}^N m(v_i - V_h)^2 g_0 w_i \delta^3(x - x_i) \end{aligned}$$

where sum is over the particles, m mass of the particle, $g_0 w_i$ is the perturbed phase density, and V_h is the hot flow velocity

- discretize to the poloidal plane by performing a Fourier transform,^a

$$\delta p_n(R, Z) = \sum_{i=1}^N m(v_i - V_h)^2 g_0 w_i \delta(R - R_i) \delta(Z - Z_i) e^{-in\phi_i}$$

^afor large n runs, a conventional ϕ deposition with a **FFT** is performed

- express lhs $\delta p_n(R, Z)$ in the finite element basis,

$$\delta p_n(R, Z) = \sum_k \delta p_n^k N^k(\eta, \xi)$$

where sum k is over the basis functions and (η, ξ) are functions of (R, Z)

- project onto the finite element space by casting in **weak form**:

$$\int N^l \sum_k \delta p_n^k N^k d^3x = \int N^l \sum_{i=1}^N m(v_i - V_h)^2 g_0 w_i \delta(R - R_i) \delta(Z - Z_i) e^{-in\phi_i} d^3x$$

$$\underline{\mathbf{M}} \delta p_n^k = \sum_{l \in k} \sum_{i=1}^N m(v_i - V_h)^2 g_0 w_i e^{-in\phi_i} N^l(\eta_i, \xi_i)$$

\mathbf{M} is the finite element mass matrix

- **gather** is done using the same shape functions

$$A(\mathbf{x}_i) = \sum_l A^l N^l(\eta_i, \xi_i)$$

for some field quantity A



The δf PIC method^{a b}

- PIC is a Lagrangian simulation of phase space $f(\mathbf{x}, \mathbf{v})$
- PIC evolves the $f(\mathbf{x}(\mathbf{t}), \mathbf{v}(\mathbf{t}))$
- spatial grid is not inherently necessary, but very convenient!
- in principle, $f(\mathbf{x}(\mathbf{t}), \mathbf{v}(\mathbf{t}))$ contains everything
- typically PIC is noisy, can't beat $1/\sqrt{N}$
- δf PIC **reduces the discrete particle noise** associated with conventional PIC
- Vlasov Equation

$$\frac{\partial f(\mathbf{z})}{\partial t} + \dot{\mathbf{z}} \cdot \frac{\partial f(\mathbf{z})}{\partial \mathbf{z}} = 0$$

\mathbf{z} is the phase coordinate

^aS. E. Parker and W. W. Lee, 'A fully nonlinear characteristic method for gyro-kinetic simulation', *Physics of Fluids B*, **5**, 1993

^bG. Hu and J. A. Krommes, "Generalized weighting scheme for δf particle simulation method", *Physics of Plasmas*, **1**, 1994



- split phase space distribution into steady state and evolving perturbation:

$$f = f_{eq}(\mathbf{z}) + \delta f(\mathbf{z}, t)$$

- δf evolves along the characteristics $\dot{\mathbf{z}}$ (control variates MC^c)

$$\delta \dot{f} = -\tilde{\mathbf{z}} \cdot \frac{\partial f_{eq}}{\partial \mathbf{z}}$$

using $\mathbf{z} = \mathbf{z}_{eq} + \tilde{\mathbf{z}}$ and $\dot{\mathbf{z}}_{eq} \cdot \frac{\partial f_{eq}}{\partial \mathbf{z}} = 0$

- the drift kinetic equations of motion are used as the particle characteristics

$$\dot{\mathbf{x}} = v_{\parallel} \hat{\mathbf{b}} + \frac{m}{eB^4} \left(u^2 + \frac{v_{\perp}^2}{2} \right) \left(\mathbf{B} \times \nabla \frac{B^2}{2} \right) + \frac{\mathbf{E} \times \mathbf{B}}{B^2} + \frac{\mu_0 m v_{\parallel}^2}{eB^2} \mathbf{J}_{\perp}$$

$$m \dot{v}_{\parallel} = -\hat{\mathbf{b}} \cdot (\mu \nabla B - e \mathbf{E})$$

^cA. Y. Aydemir, "A unified MC interpretation of particle simulations...", *Physics of Plasmas*, **1**, 1994

Slowing Down Distribution for Hot Particles

- for the slowing down distribution function

$$f_{eq} = \frac{P_0 \exp\left(\frac{P_\zeta}{\psi_0}\right)}{\varepsilon^{3/2} + \varepsilon_0^{3/2}}$$

where $P_\zeta = g\rho_{\parallel} - \psi$ is the canonical toroidal momentum and ε is the energy, ψ_0 is the total flux, and ε_c is the critical slowing down energy

$$\begin{aligned} \dot{\delta f} = & f_{eq} \left\{ \frac{mg}{e\psi_0 B^3} \left[\left(v_{\parallel}^2 + \frac{v_{\perp}^2}{2} \right) \delta \mathbf{B} \cdot \nabla B - \mu_0 v_{\parallel} \mathbf{J} \cdot \mathbf{E} \right] \right. \\ & \left. + \frac{\delta \mathbf{v} \cdot \nabla \psi_p}{\psi_0} + \frac{3}{2} \frac{e\varepsilon^{1/2}}{\varepsilon^{3/2} + \varepsilon_0^{3/2}} \mathbf{v}_D \cdot \mathbf{E} \right\} \end{aligned}$$

where

$$\begin{aligned} \mathbf{v}_D &= \frac{m}{eB^3} \left(v_{\parallel}^2 + \frac{v_{\perp}^2}{2} \right) (\mathbf{B} \times \nabla B) + \frac{\mu_0 m v_{\parallel}^2}{eB^2} \mathbf{J}_{\perp} \\ \delta \mathbf{v} &= \frac{\mathbf{E} \times \mathbf{B}}{B^2} + \mathbf{v}_{\parallel} \cdot \frac{\delta \mathbf{B}}{B} \end{aligned}$$

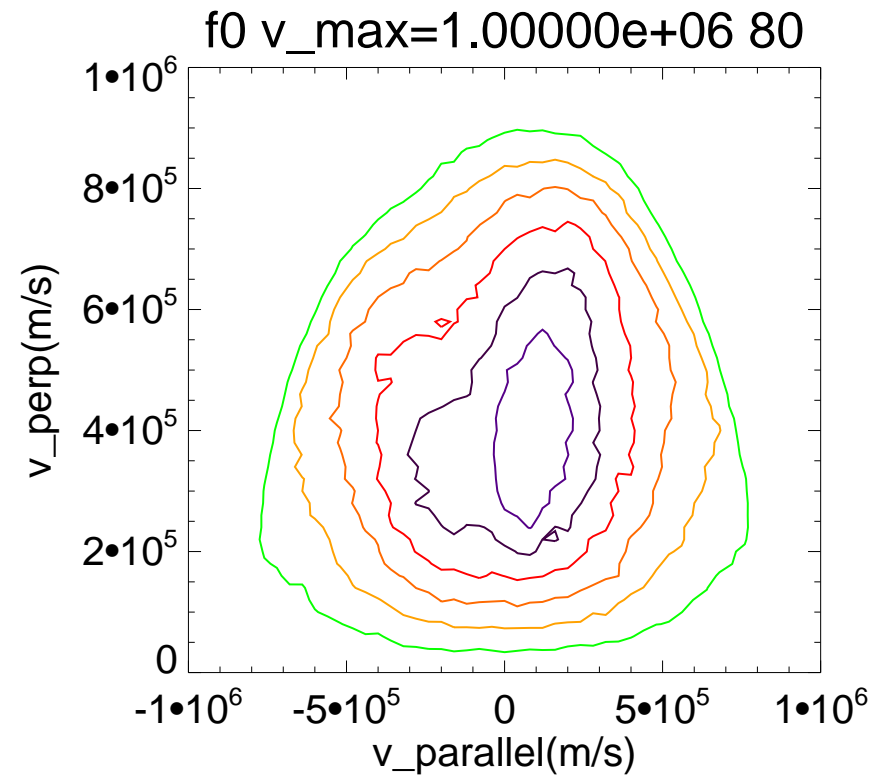
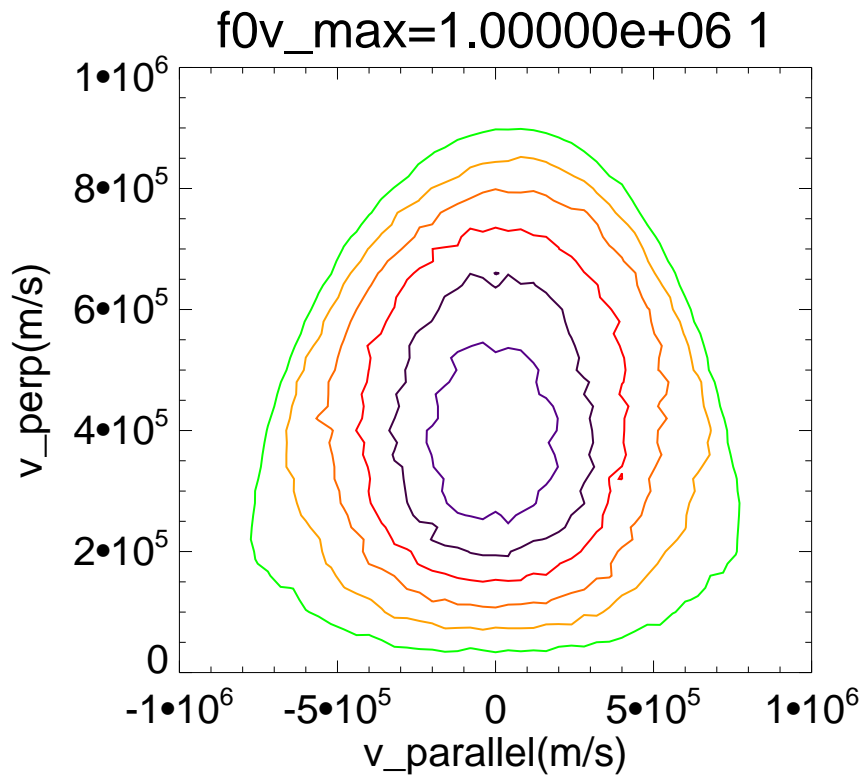
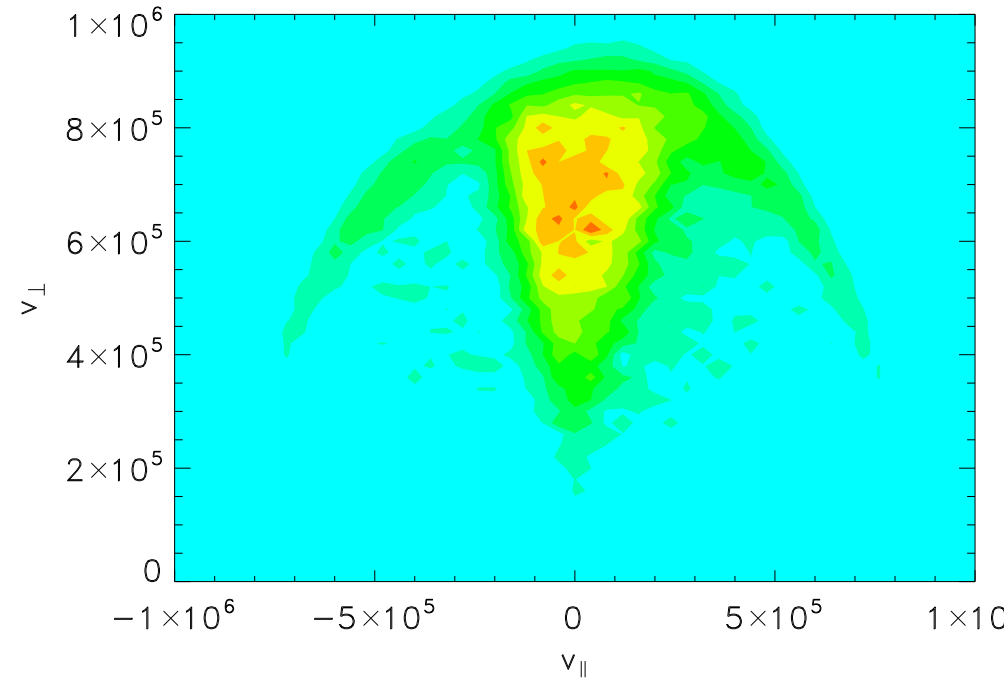
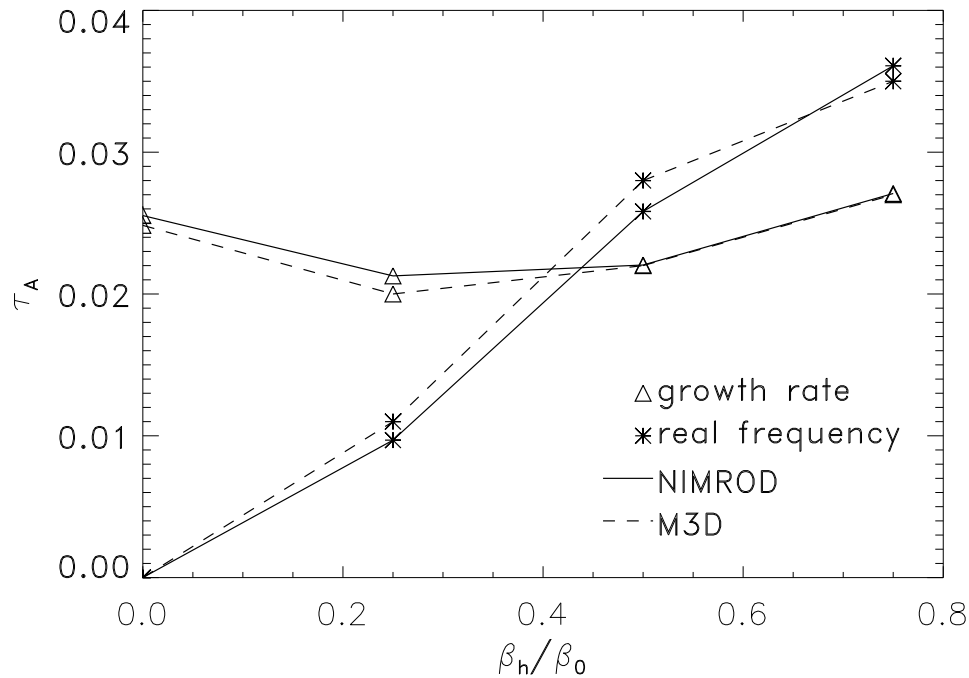


Figure 1: f_0 at $v_{max} = 1.0 \times 10^6$ m/s initial load and steady state

Benchmark with M3D



- ideal (1, 1) internal kink mode $\beta = 8\%$
- velocity shows most activity in the most energetic particles
- mostly in trapped region, also in extremes of passing particles

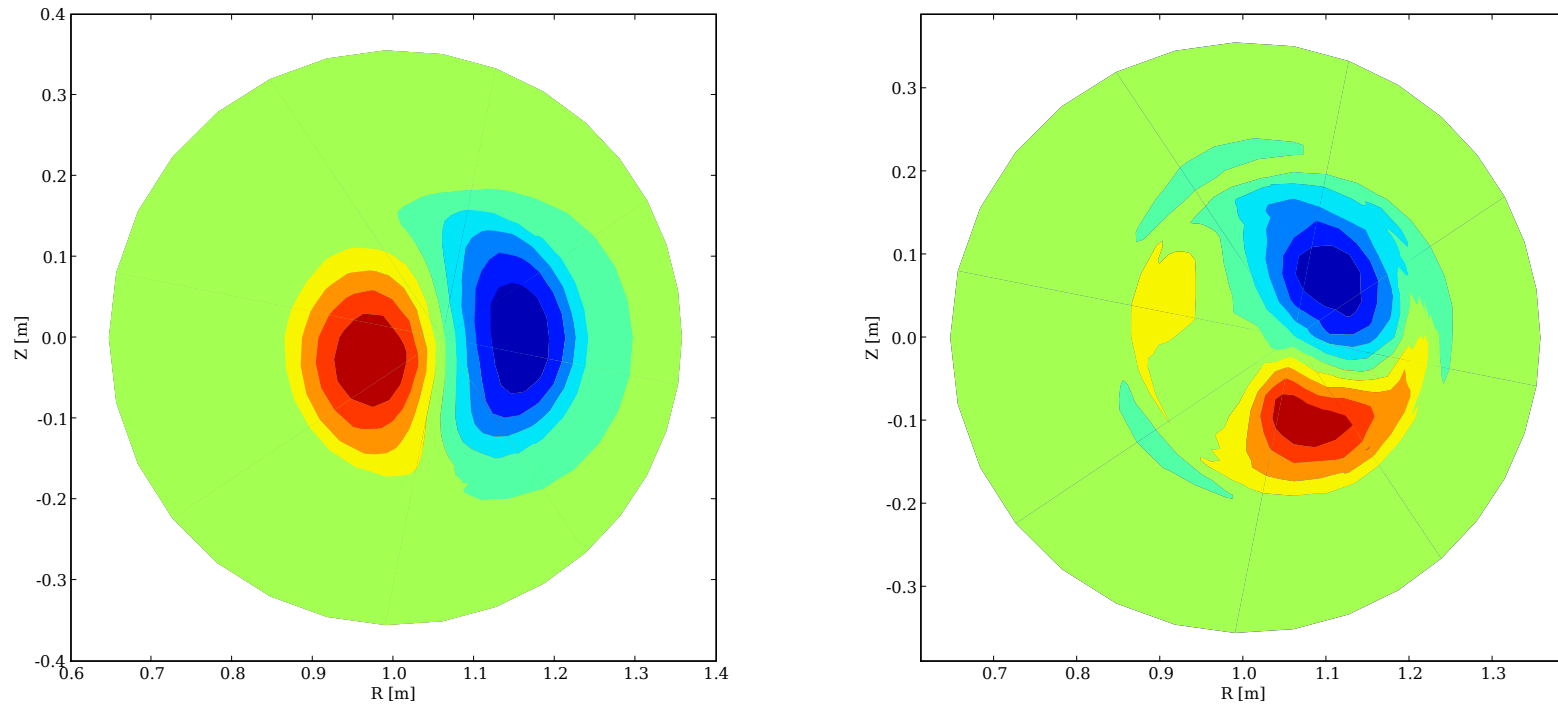


Figure 2: $n = 1$ $p_{h\perp}, \Delta p_h$

- simulations without anisotropy reproduce ideal MHD with γ within 10%
- no real frequency
- anisotropic pressure $\Delta p = p_{\parallel} - p_{\perp}$ is key to energetic effects

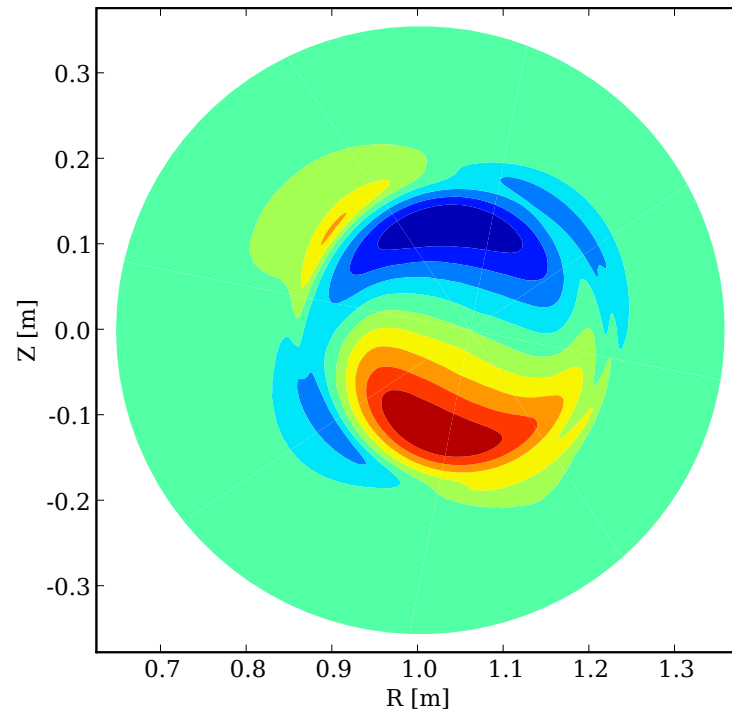
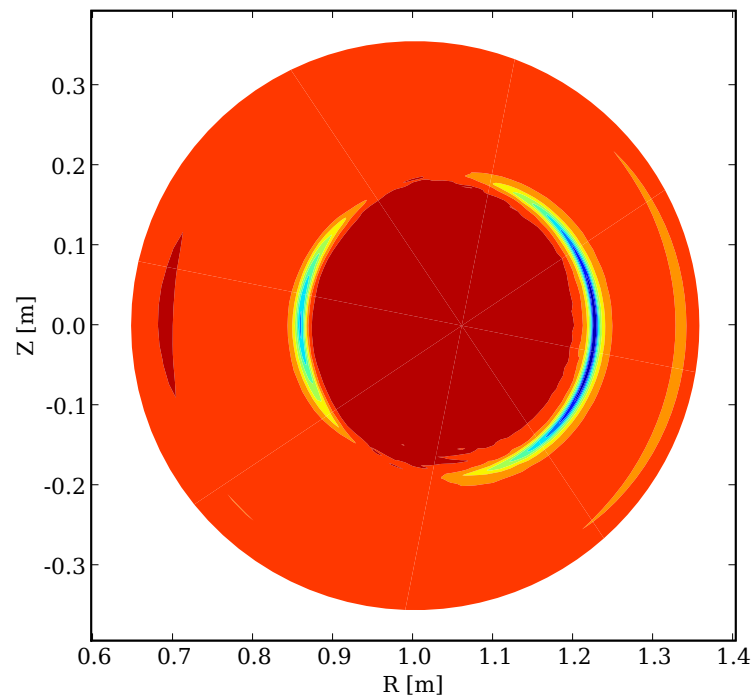


Figure 3: $n = 1V_\phi$ with passing particles, V_ϕ with trapped particles

- simulations with only passing particles stabilize but do not change V_ϕ mode topology
- trapped particles excite precessional fishbone mode (plot on right)

Impact of increasing velocity cutoff

- benchmark (1, 1) performed with $v_{max} = 1 \times 10^6 m/s$
- increasing the velocity cutoff results in stronger stabilization
 - fixed β_h - increase in energy range, decrease in density → **fewer particles are doing more**
 - increase in real frequency

– examine $\beta_h = .5$

v_{max}	γT_A	ωT_A
1.0	0.0218	0.0257
1.1	0.0180	0.0290
1.2	0.0142	0.0317
1.3	0.0093	0.0345
1.5	small	0.0370

Examination of δf in phase space

- more energetic particles drive higher frequency
- plots of velocity space density of $n = 1$ projection of δf shows dominance of trapped particles as cut-off is increased
- this is reflected in decreasing relative amplitude of passing particle “wing”
- note the asymmetry in the profile
 - trapped cone region biased toward counter propagating (negative v_{\parallel})
 - passing region biased towards co-propagating



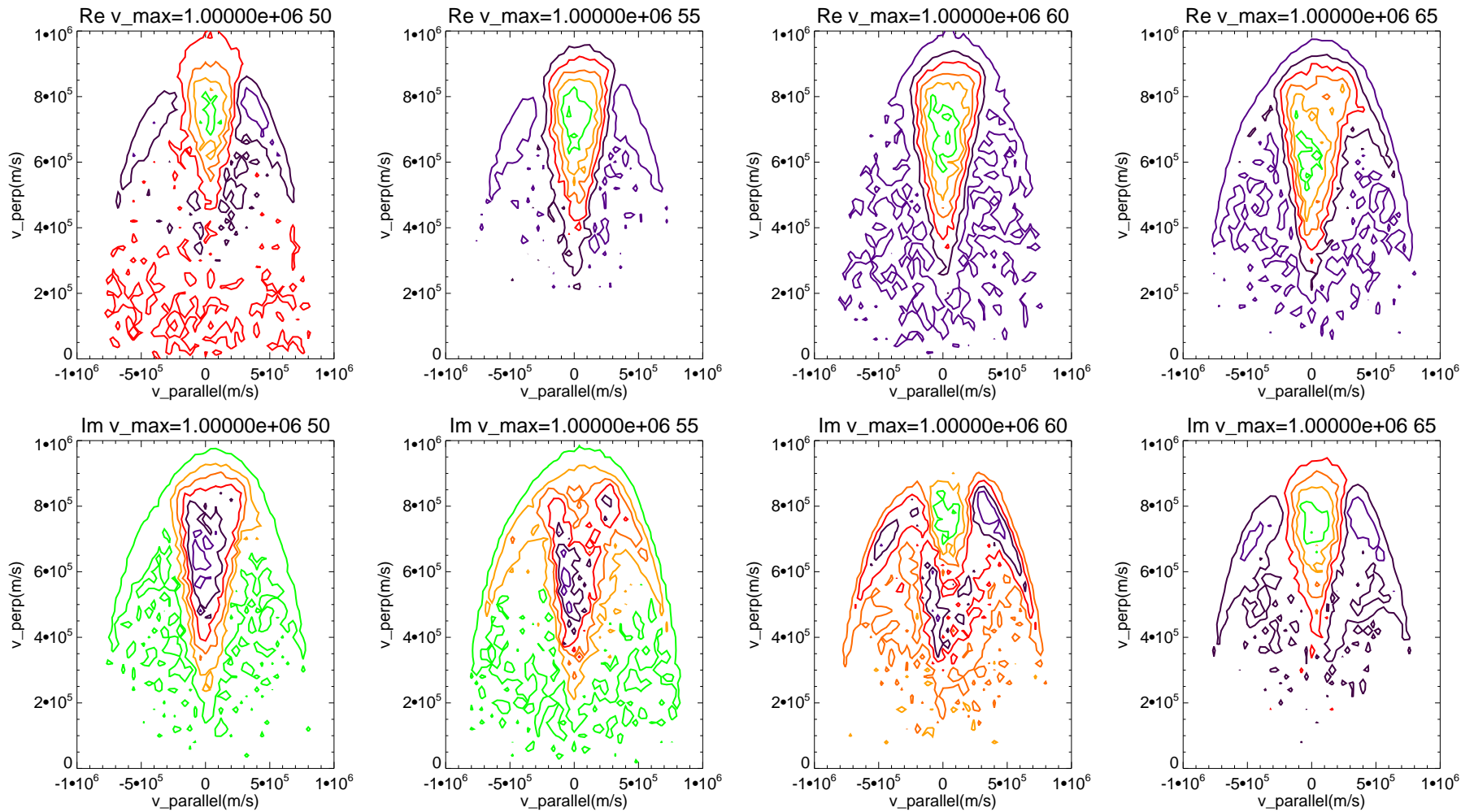


Figure 4: $n = 1$ projection of δf at $v_{max} = 1.0 \times 10^6$ m/s

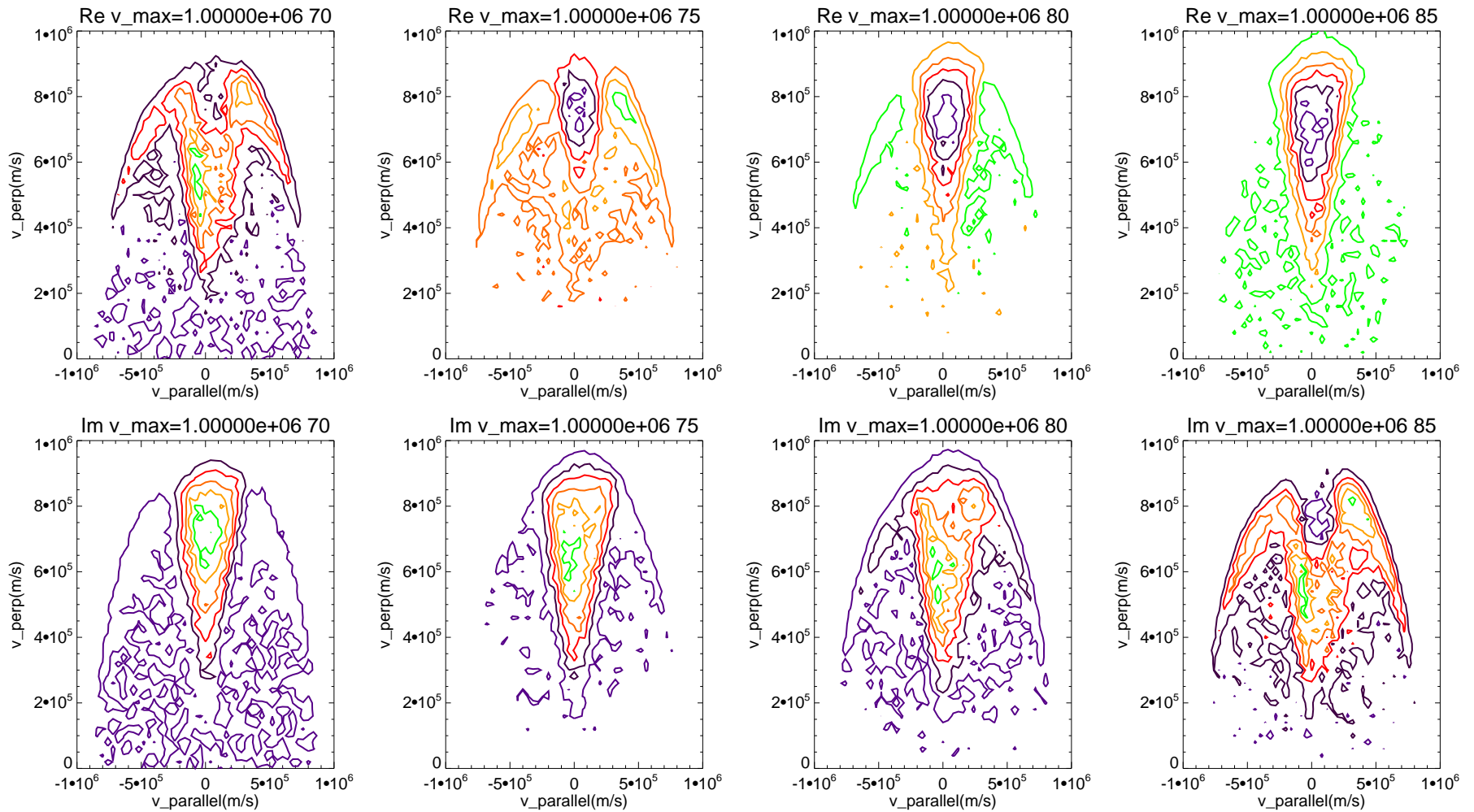


Figure 5: $n = 1$ projection of δf at $v_{max} = 1.0 \times 10^6$ m/s

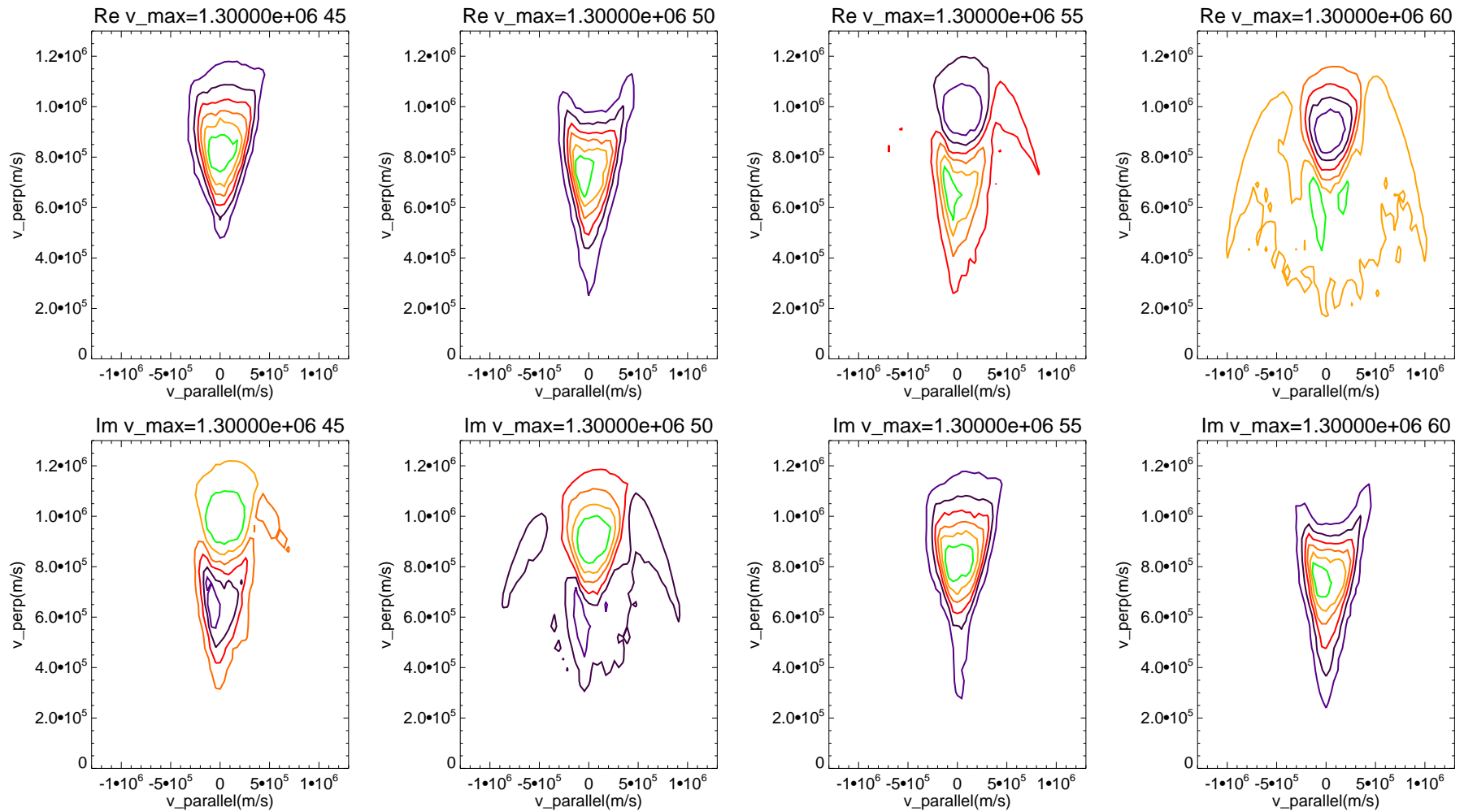


Figure 6: $n = 1$ projection of δf at $v_{max} = 1.3 \times 10^6$ m/s

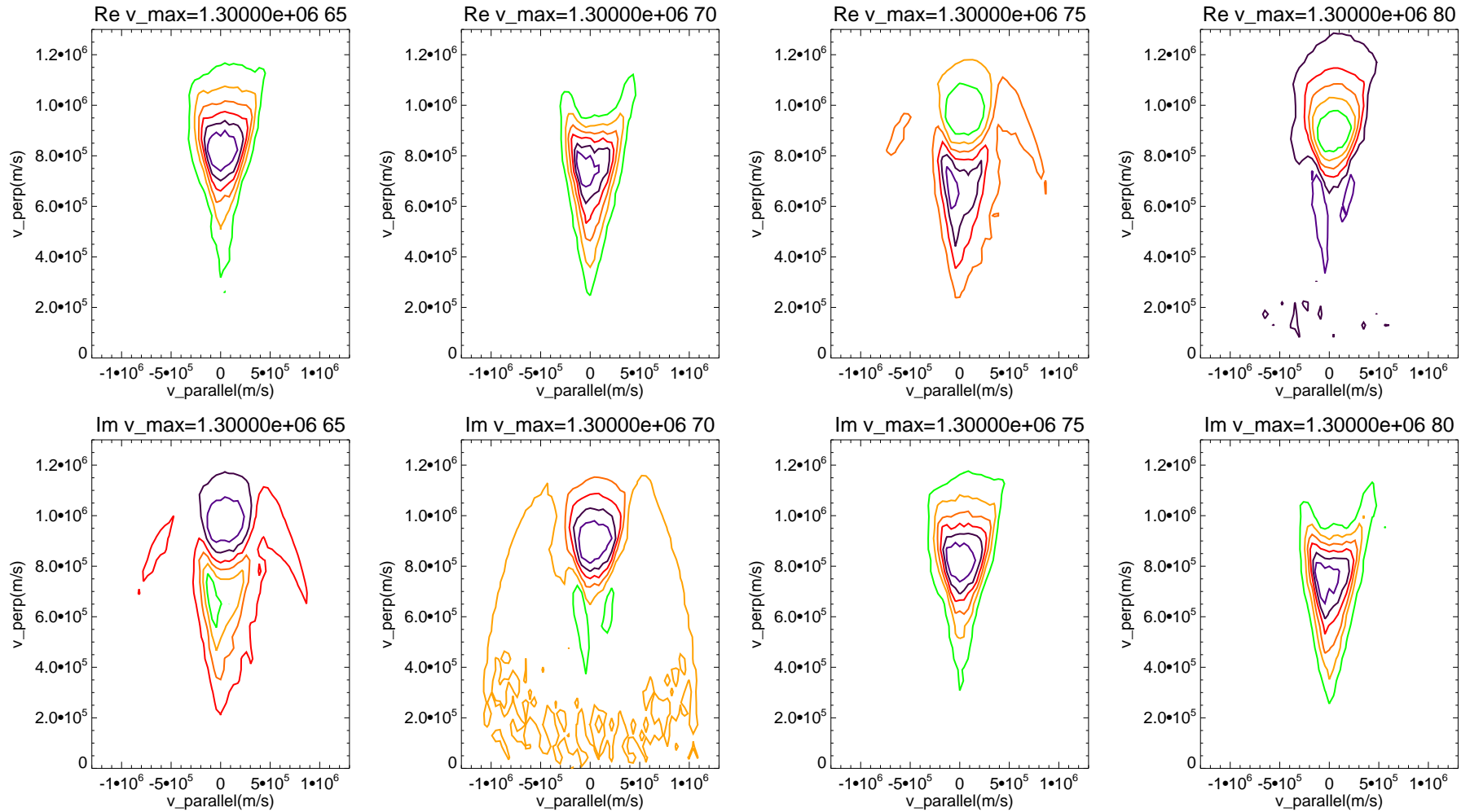


Figure 7: $n = 1$ projection of δf at $v_{max} = 1.3 \times 10^6$ m/s

Linear Simulations of Tearing Modes in a RFP

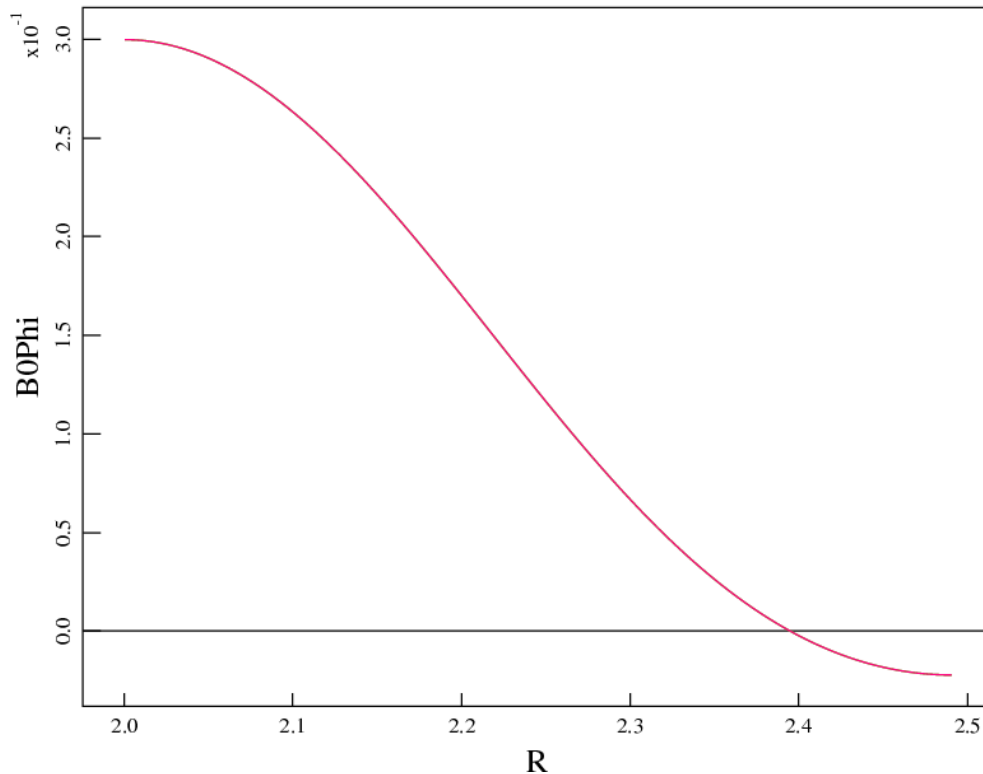
- alpha model equilibrium $\nabla \times \mathbf{B} = \mu \mathbf{B}$ $\mu = 2\Theta \left[1 - \left(\frac{r}{a} \right)^{\alpha_0} \right]$

- parameters for straight cylinder

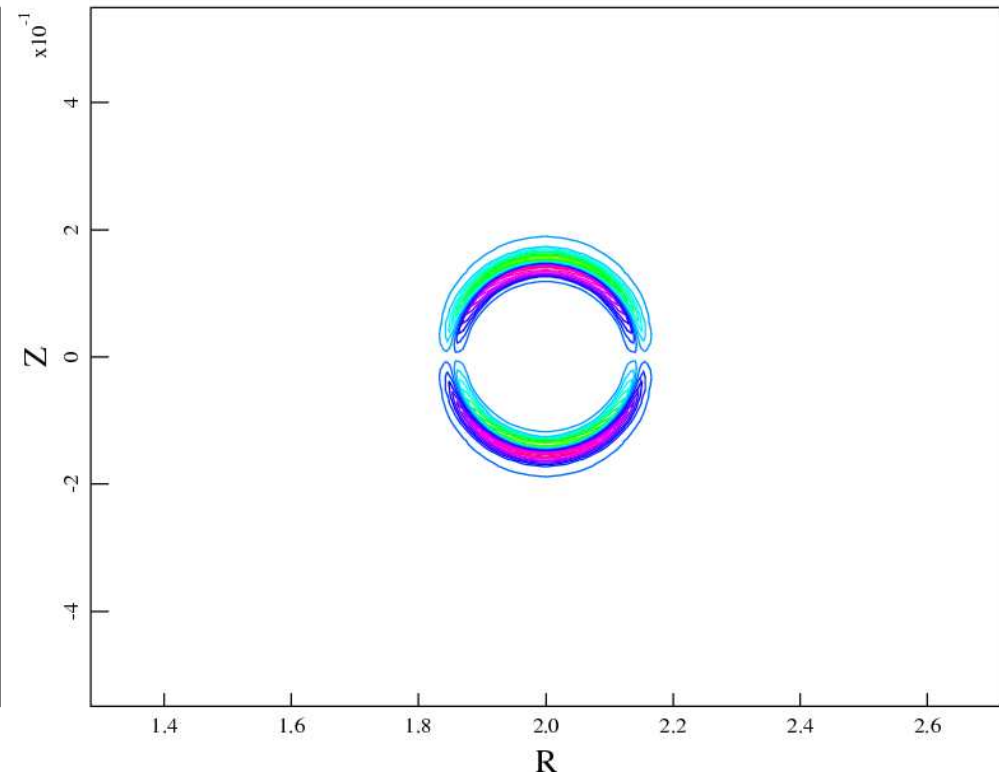
$$a = .5m, B_0 = .3T, \Theta = 1.75, \alpha_0 = 3,$$

$$S = 1.e4, ka = 2, \gamma\tau_A = 1.3e - 3$$

B0_Phi vs. R



Re VPhi



δf and the Lorentz Equations

- Lorentz equation of motion

$$\begin{aligned}\dot{\mathbf{x}} &= \mathbf{v} \\ \dot{\mathbf{v}} &= \frac{q}{m} (\mathbf{E} + \mathbf{v} \times \mathbf{B})\end{aligned}$$

- for full kinetic equations use^a

$$f_0 = f(\mathbf{x}, v^2) + \frac{1}{\omega_c} (\mathbf{v} \cdot \mathbf{b} \times \nabla f)$$

- weight equation is

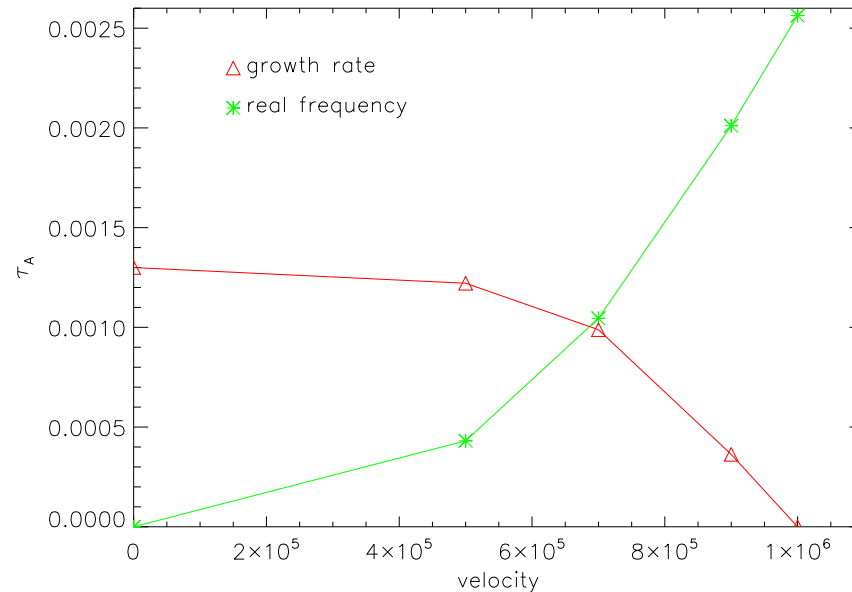
$$\dot{f} = -\frac{\delta \mathbf{E} + \mathbf{v} \times \delta \mathbf{B}}{B} \cdot \mathbf{b} \times \nabla f - \frac{2q}{m} \delta \mathbf{E} \cdot \mathbf{v} \frac{\partial f}{\partial v^2}$$

- satisfying in that it is very similar to drift kinetic weight equation

^aM. N. Rosenbluth and N. Rostoker “Theoretical Structure of Plasma Equations”, Physics of Fluids **2** 23 (1959)

FLR Stabilization of RFP Tearing Mode

- initialize with monoenergetic particles, only $\mathbf{v} \times \delta\mathbf{B}$ in weight equation
- use **only** perpendicular pressure for comparison with theory



- stabilization at $\rho_h \simeq 4cm$
- simulation sees real frequency - probably due to finite spread in velocity
- simplify problem to slab tearing mode - work is beginning

Some stuff at PSI Center

- mimetic GS solver for triangles and quads (MatLab)
 - generates equilibria suitable for NIMROD
 - uses NIMROD finite element nodes
 - good equilibria for HIT-SI, FRCs, (LDX soon)
 - see C. Akcay poster 2C-6 Tue 4:30-6:30
- migrating to NIMDevel (would like to set up mirror)
 - particles are running
 - porting experiments in process
- LDX simulations show promise - show movie
- optimization flags - turn off IEEE conformance



Future Work

- add more diagnostics
 - look at spatial dependence of phase space
 - synthetic diagnostics
- make particle domain decomp independent of fluids
- high order particles!
- apply to PSI Center experiments
- continue collaboration with MST
 - tentative extended visit in early June

