

# ITG-driven instabilities in extended MHD

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# Can ion-temperature-gradient (ITG) instabilities exist in nimrod simulations?

- ▶ ITG effects found in recent nimrod simulations of DIII-D experiments [Schnack 10]
- ▶ Extended MHD model with two-fluid physics implemented in NIMROD code [Sovinec *et al.* 04].
- ▶ To what extent are ITG physics captured in extended MHD nimrod simulations?
  - ▶ May help explain unstable modes in MHD-stable configurations.
  - ▶ May help distinguish truly numerical instabilities.
- ▶ Recent theory and nimrod calculations developed.

# Single fluid formulation of extended MHD model is implemented in NIMROD code [Sovinec *et al.*, 2004]

$$\frac{d\rho}{dt} = -\rho \nabla \cdot \mathbf{u} + D \nabla^2 \rho \quad (1)$$

$$\rho \frac{d\mathbf{u}}{dt} = -\nabla p + \mathbf{J} \times \mathbf{B} - \nabla \cdot \boldsymbol{\pi} \quad (2)$$

$$\frac{n}{\gamma - 1} \frac{dT}{dt} = -\frac{p}{2} \nabla \cdot \mathbf{u} - \boldsymbol{\pi} : \nabla \mathbf{u} - \nabla \cdot \mathbf{q} + Q \quad (3)$$

$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E} \quad (4)$$

$$\mu_0 \mathbf{J} = \nabla \times \mathbf{B} \quad (5)$$

$$\mathbf{E} = -\mathbf{u} \times \mathbf{B} + \eta \mathbf{J} + \frac{\lambda}{ne} (\mathbf{J} \times \mathbf{B} - \nabla(1 - \tau)p) \quad (6)$$

where

$$\boldsymbol{\pi}_i = \frac{\tau p}{4\Omega} \left\{ [\mathbf{b} \times (\nabla \mathbf{u} + \nabla \mathbf{u}^T) \cdot (\mathbf{I} + 3\mathbf{b}\mathbf{b})] + [\text{previous term}]^T \right\} \quad (7)$$

# Two-fluid equilibrium in shear-less slab configuration and flute-like perturbation are assumed

## ► Equilibrium

$$\mathbf{u} = 0 \quad (8)$$

$$\mathbf{B} = B\mathbf{e}_z \quad (9)$$

$$\frac{d}{dx} \left( p + \frac{B^2}{2} \right) = \rho \mathbf{g} \cdot \mathbf{e}_x = \rho g \quad (10)$$

$$n\mathbf{E} = \nabla p_i - \rho \mathbf{g}. \quad (11)$$

## ► Perturbation

$$\mathbf{u} = (u_x(x), u_y(x), u_z(x)) \exp(-i\omega t + ik_y y + ik_z z). \quad (12)$$

where  $k_z/k_y \sim \epsilon$ ,  $k_y L_A \sim \epsilon^{-1}$ ,  $k_y d_i \sim \lambda \sim \delta$ ,  $u_y \sim \epsilon u_x$ ,  $u_x \sim u_z$ ,  $\epsilon \ll 1$ , where  $L_A = |d \ln A / dx|^{-1}$  is the spatial scale of field  $A$  in  $x$  direction, and  $d_i = \sqrt{\rho_i / \rho} / \Omega$  is the ion Larmor radius.

# Ideal MHD model yields coupling of shear-Alfvén and slow wave modes

- ▶ In absence of  $\mathbf{g}$

$$(\omega^2 - k_z^2 u_A^2) \left( \omega^2 - \frac{k_z^2 c_s^2}{1 + \gamma\beta} \right) = 0 \quad (13)$$

- ▶ 3D  $g$  mode

$$\omega^4 + \left[ \frac{g^2}{u_A^2 (1 + \gamma\beta)} - \frac{\rho'}{\rho} g - k_z^2 u_A^2 \frac{1 + 2\gamma\beta}{1 + \gamma\beta} \right] \omega^2 + \frac{k_z^2 u_A^2}{1 + \gamma\beta} \left( k_z^2 c_s^2 - \frac{g^2}{u_A^2} + \gamma\beta g \frac{\rho'}{\rho} \right) = 0 \quad (14)$$

# Extended MHD model leads to linear dispersion relation with ITG effects [Zhu et al. 2011]

To dominant order in  $\epsilon$ , the dispersion relation takes the quintic form

$$c_5 \omega^5 + c_4 \omega^4 + c_3 \omega^3 + c_2 \omega^2 + c_1 \omega + c_0 = 0 \quad (15)$$

where

$$c_5 = 1 + \gamma\beta + \frac{k_y^2 \delta^2 \tau^2}{4\Omega^2} \frac{\rho}{\rho'} \beta \quad (16)$$

$$c_4 = -\frac{k_y^3 \tau^3}{4\Omega^3} \frac{\rho \rho'}{\rho^2} \beta(1 + \beta) + \frac{k_y \lambda}{\Omega} \left[ (1 + \tau + \gamma\beta) \frac{\rho'}{\rho} - \tau c_s^2 \frac{\rho'}{\rho} \right] + \dots \quad (17)$$

$$c_3 = -k_z^2 u_A^2 (1 + 2\gamma\beta) - \frac{k_y^2 \lambda^2}{\Omega^2} \left[ k_z^2 u_A^2 c_s^2 + \tau \frac{\rho'}{\rho} \left( -\frac{\rho'}{\rho} + c_s^2 \frac{\rho'}{\rho} \right) \right] + \dots \quad (18)$$

$$c_2 = -\frac{k_y \lambda}{\Omega} k_z^2 u_A^2 \left[ (1 + \tau + \gamma\beta) \frac{\rho'}{\rho} - \tau c_s^2 \frac{\rho'}{\rho} \right] + \frac{k_y \delta \tau}{\Omega} k_z^2 u_A^2 (1 + \beta) \frac{\rho'}{\rho} + \dots \quad (19)$$

$$c_1 = k_z^4 u_A^2 c_s^2 + \frac{k_y^2 \lambda^2}{\Omega^2} k_z^2 u_A^2 \tau \frac{\rho'}{\rho} \left( -\frac{\rho'}{\rho} + c_s^2 \frac{\rho'}{\rho} \right) + \dots \quad (20)$$

$$c_0 = \frac{k_y^3 \lambda^2 \delta \tau}{\Omega^3} k_z^2 u_A^2 (1 + \beta) \left( \frac{\rho'}{\rho} - c_s^2 \frac{\rho'}{\rho} \right) \left( \frac{\rho'}{\rho} \right)^2 \quad (21)$$

## Further reduction can be made in low frequency and small Larmor radius regime

- ▶ To first order in  $(\omega_{*i}, \omega_{*p}) \ll \omega_A$  and  $k_y \rho_L \ll 1$ , the dispersion can be further reduced to

$$(1 + \gamma\beta) \left(\frac{\omega}{\omega_A}\right)^4 - \lambda \left[ \frac{\omega_{*i}}{\omega_A} \left(\eta - \frac{2}{3}\right) + \frac{\omega_{*p}}{\omega_A} (1 + \gamma\beta) \right] \left(\frac{\omega}{\omega_A}\right)^3 - (1 + 2\gamma\beta) \left(\frac{\omega}{\omega_A}\right)^2 + \left\{ \lambda \frac{\omega_{*i}}{\omega_A} \left(\eta - \frac{2}{3}\right) + \frac{\omega_{*p}}{\omega_A} [\lambda(1 + \gamma\beta) - \delta\tau(1 + \beta)] \right\} \left(\frac{\omega}{\omega_A}\right) + \gamma\beta = 0 \quad (22)$$

where  $\eta = d \ln T / d \ln \rho$ ,  $T = p/\rho$ ,  $\gamma = 5/3$ , and

$$\omega_{*i} = -\frac{k_y \tau p \rho'}{\Omega \rho \rho}, \quad \omega_{*p} = -\frac{k_y p \rho'}{\Omega \rho \rho} \quad (23)$$

- ▶ Formally, the zero  $\beta$  limit gives the cubic form of dispersion relation

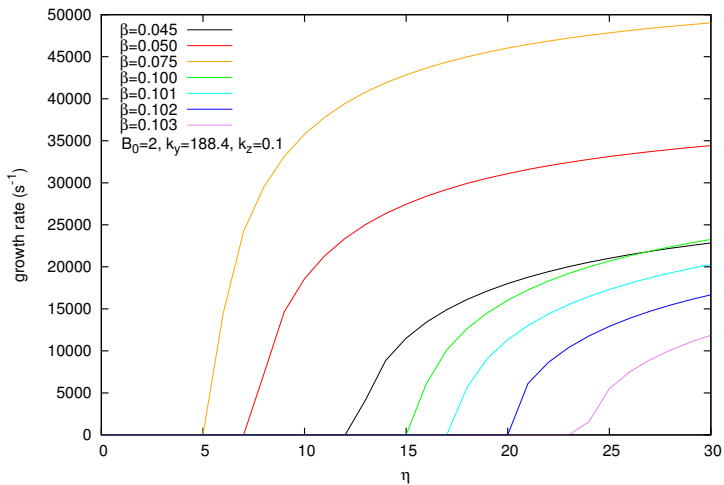
$$\left(\frac{\omega}{\omega_A}\right)^3 - \lambda \frac{\omega_{*i}}{\omega_A} \left(\eta - \frac{2}{3} + \frac{\eta + 1}{\tau}\right) \left(\frac{\omega}{\omega_A}\right)^2 - \frac{\omega}{\omega_A} + \lambda \frac{\omega_{*i}}{\omega_A} \left(\eta - \frac{2}{3}\right) + \frac{\omega_{*p}}{\omega_A} (\lambda - \delta\tau) = 0 \quad (24)$$

# Two cases are considered for the solution of dispersion relation

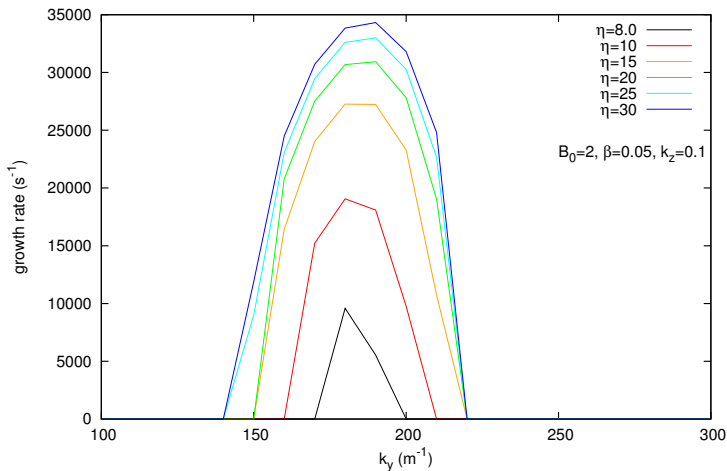
- ▶ Case 1 [Schnack *et al.* 11]:  $B_0 = 2$ ,  $\beta = 0.05$ ,  $\tau = 0.2$
- ▶ Case 2 [Zhu *et al.* 07]:  $B_0 = 6$ ,  $\beta = 0.1$ ,  $\tau = 0.5$
- ▶ Parameters common to both cases:  
 $\rho = 2 \times 10^{20} m_i$ , density (pressure) gradient scale length  $L_\rho$  ( $L_p$ ).



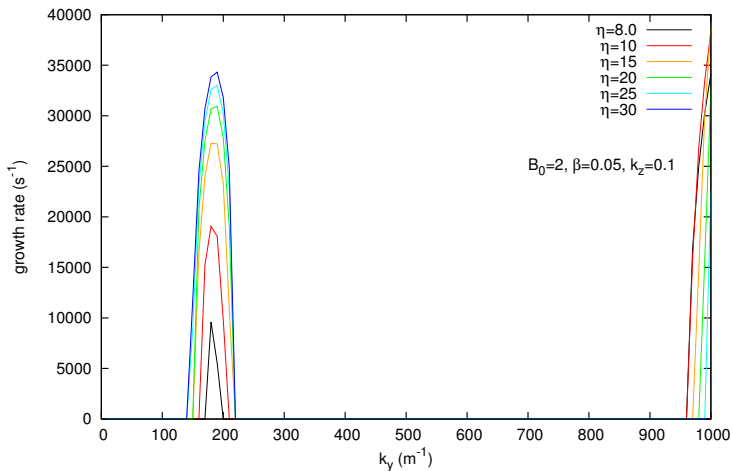
# $\eta$ thresholds for ITG-driven instability onset are sensitive to $\beta$



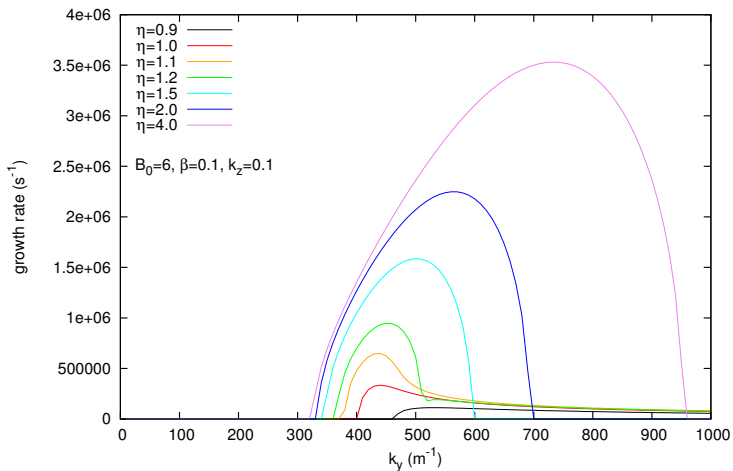
# Solutions from analytic dispersion relation show instabilities when $\eta > 8$ in $k_y \rho_L \sim 0.1$ regime



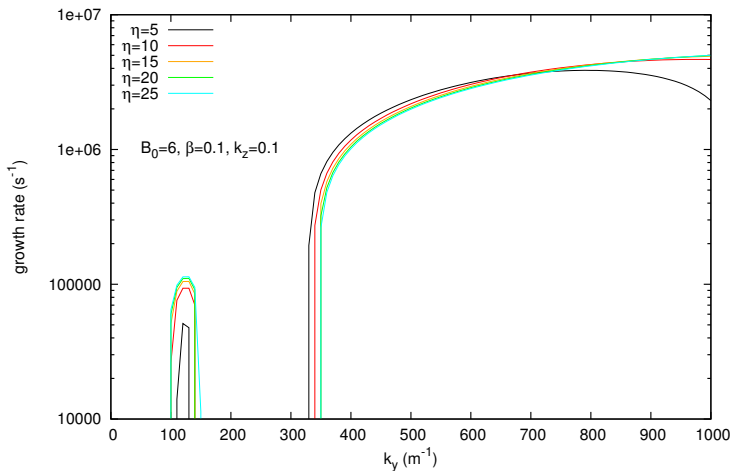
# In $k_y \rho_L \sim 0.1 - 1$ regime there are two branches of ITG-driven instabilities



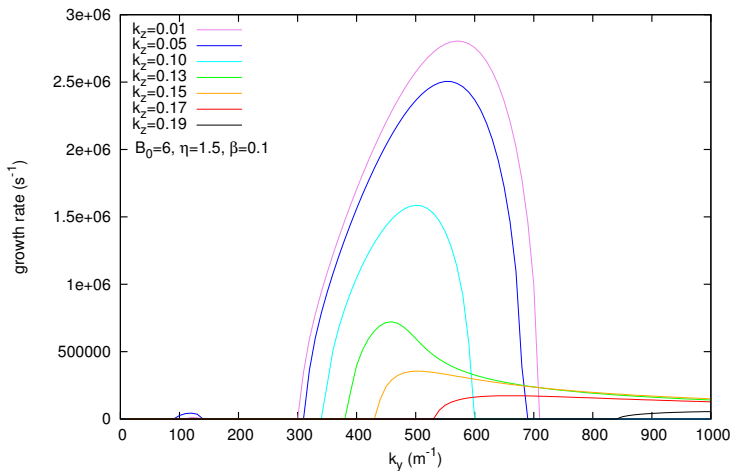
# High ITG branch ( $k_y \rho_L \sim 1$ ) more dominant in higher $\beta$ regime ( $\eta_{\text{crit}} \sim 0.8$ )



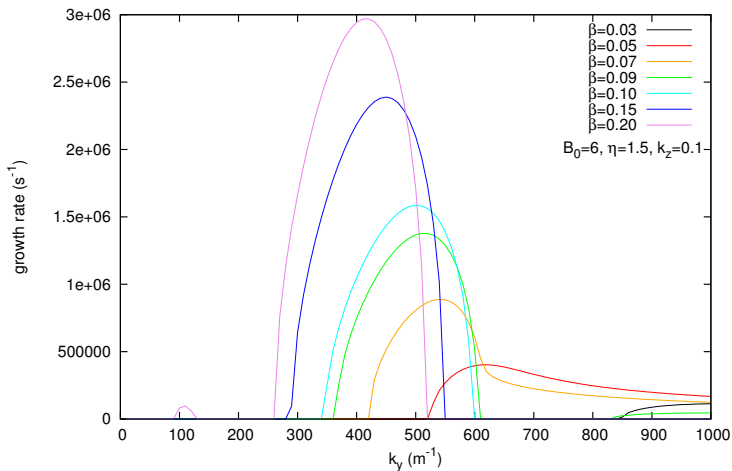
# Low ITG branch ( $k_y \rho_L \sim 0.1$ ) becomes less dominant in higher $\beta$ regime



Instabilities more unstable at smaller  $k_z$  and have a cut-off  $k_z \sim 0.2$  in high branch ( $k_y \rho_L \sim 1$ )



High branch ( $k_y \rho_L \sim 1$ ) growth increases with  $\beta$  ( $\beta_{\text{crit}} \sim 0.025$ ) and shifts towards smaller  $k_y \rho_L$

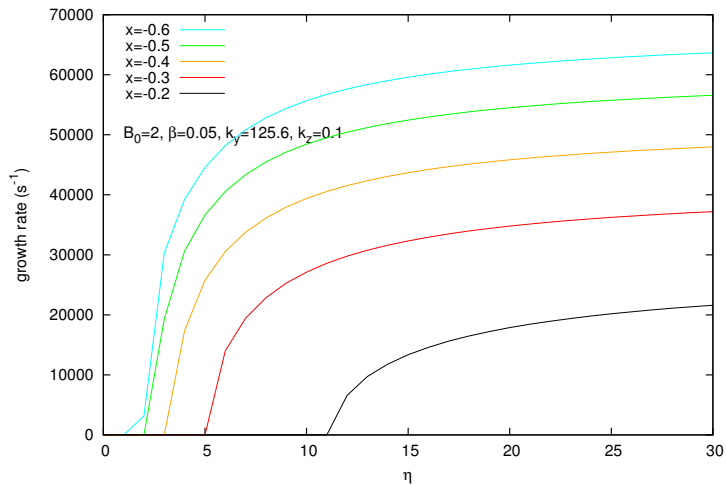


# Exponential profiles for both density and pressure are considered in nimrod calculation

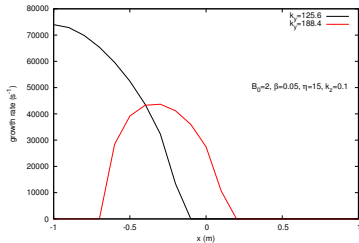
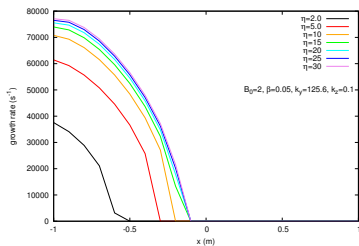
- ▶ Density and pressure:  $\rho(x) = \rho_0 e^{-\frac{x}{L_\rho}}$ ,  $p(x) = p_0 e^{-\frac{x}{L_p}}$ ,  
 $\eta = \frac{L_\rho}{L_p} - 1$
- ▶ Magnetic field  $\mathbf{B} = B(x)\hat{\mathbf{Z}}$  is in  $R - Z$  plane.
- ▶  $k_y$  of perturbation is in periodic direction.
- ▶ While  $\eta$  is constant, other key parameters such as  $\beta$  can have large range of variation over  $x$ .



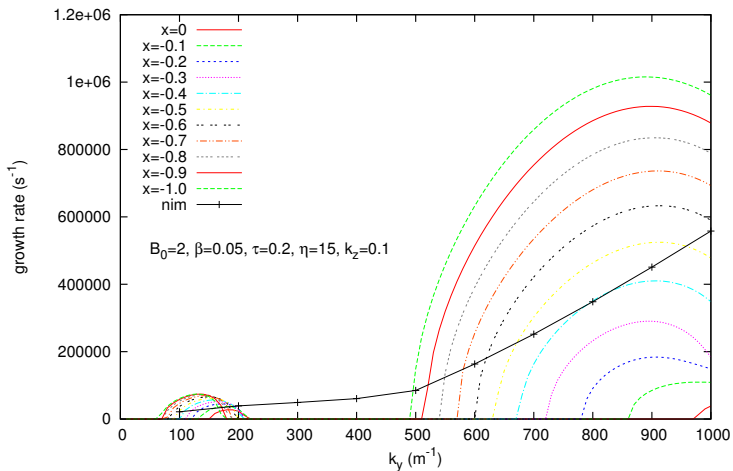
# For a fixed $k_y$ ITG growth at each location in $x$ has its own $\eta$ dependence



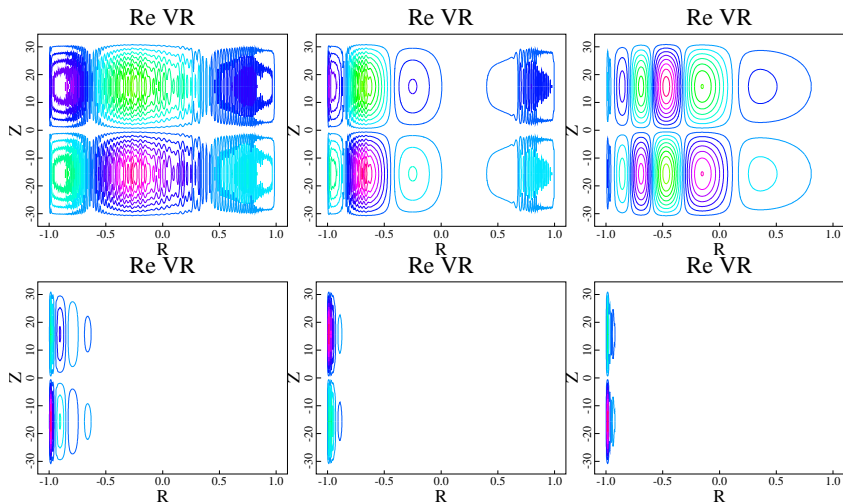
For a fixed  $\eta$  or  $k_y$  each location in  $x$  is associated with an ITG growth from local dispersion relation



# Comparison between nimrod calculation and theory show qualitative consistency

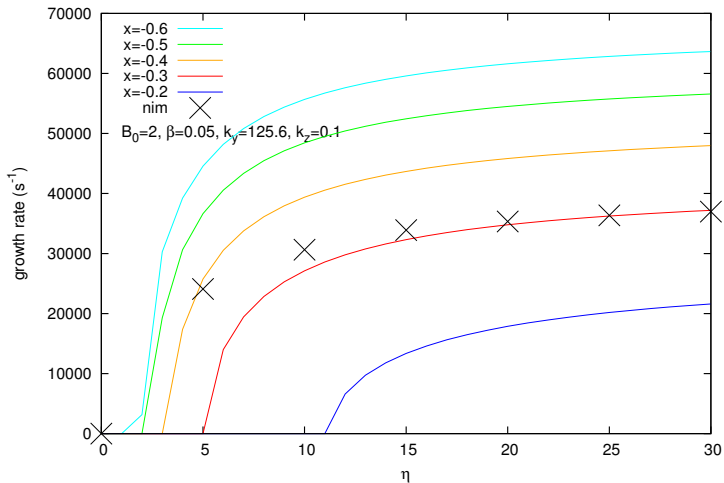


As  $k_y$  increases mode structure becomes more localized on the most unstable flux surface in  $x$



Upper row:  $k_y = 100, 200, 400$ ; Lower row:  $k_y = 600, 800, 1000$

# Comparison between nimrod calculation and theory show qualitative agreement when spatial profile effects taken into account



# Summary

- ▶ Recent theory and nimrod calculations have identified ITG-driven instabilities in an extended MHD model.
- ▶ Extended MHD theory based on single-fluid formulation is able to explain observations from nimrod calculations
  - ▶ Two branches of ITG-driven growth in  $k_y$  space
  - ▶ Asymmetric spatial structure of ITG-mode in exponentially shaped equilibrium
- ▶ Discrepancy may due to spatial profile effects.
- ▶ Future work
  - ▶ Comparison with two-fluid formulation.
  - ▶ NIMROD calculation in other equilibrium profiles better for comparison with local dispersion relation.