

Numerical Magnetohydrodynamics and Extended MHD for Magnetic Confinement

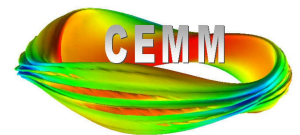
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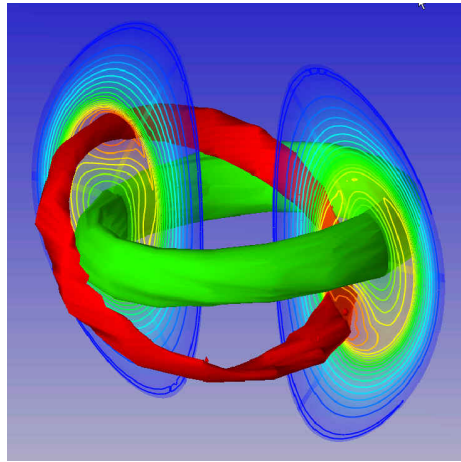
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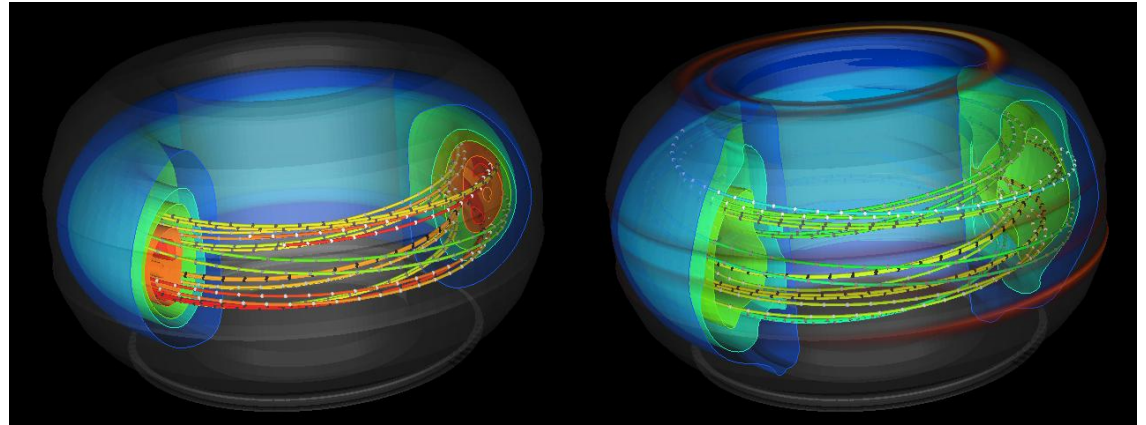
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Introduction: Magnetic confinement systems are rich in macroscopic dynamics, and ...



Simulation of internal kink in NSTX by W. Park, PPPL.

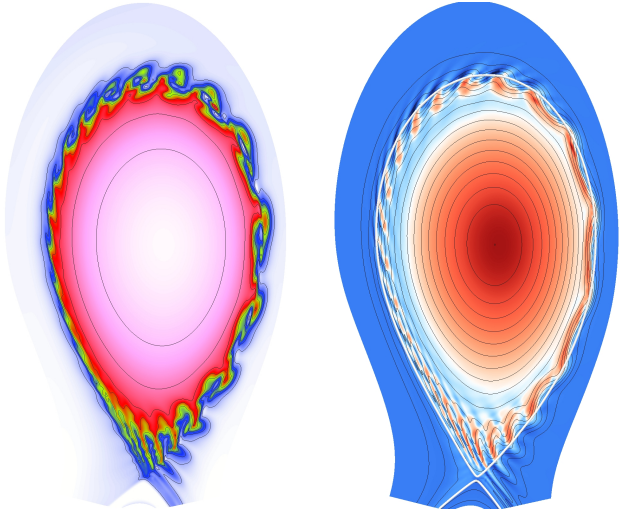


High-pressure disruption simulation by S. Kruger and A. Sanderson [Phys. Plasmas **12**, 56113].

- Tokamak sawteeth
 - Magnetic reconnection
 - Energetic-particle effects
- Tokamak disruption
 - Multi-physics effects in different forms of disruption
 - Mitigation systems

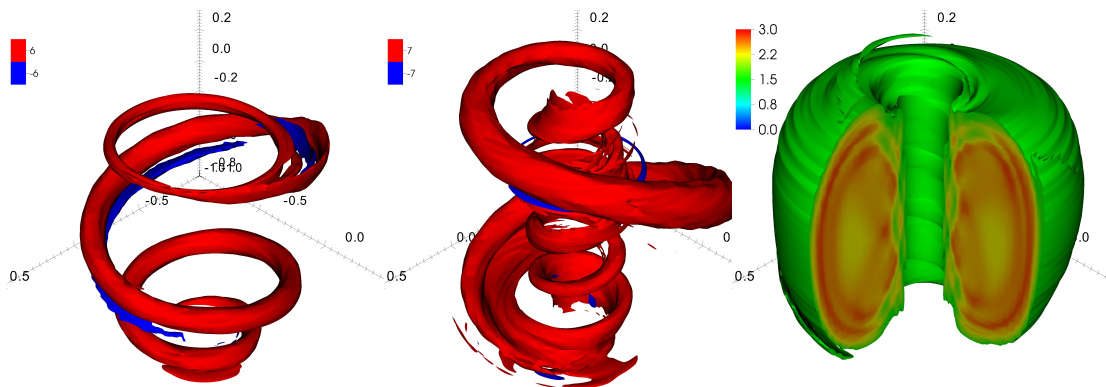


... numerical simulation of macroscopic dynamics provides important information.



Particle and current density from an ELM simulation by G. Huysmans [PPCF 51, 124012].

- Edge-localized modes (ELMs)
 - Mode coupling
 - Resonant perturbation effects
- Magnetic relaxation
 - Magnetic island evolution
 - Dynamo effects in RFPs and spheromaks
 - Non-inductive current drive



Simulation of Pegasus startup by J. O'Bryan [PPCF 56, 064005].



Outline

- Introduction
- Background
 - Physical parameters
 - Confinement and sources of free energy
 - Resonances
- Pertinent examples
- Models for macroscopic dynamics
 - Primitive-variable systems
 - Reduced systems
- Numerical methods
 - Time-advance
 - Spatial representation
- Parallel computing
- Open challenges and outlook



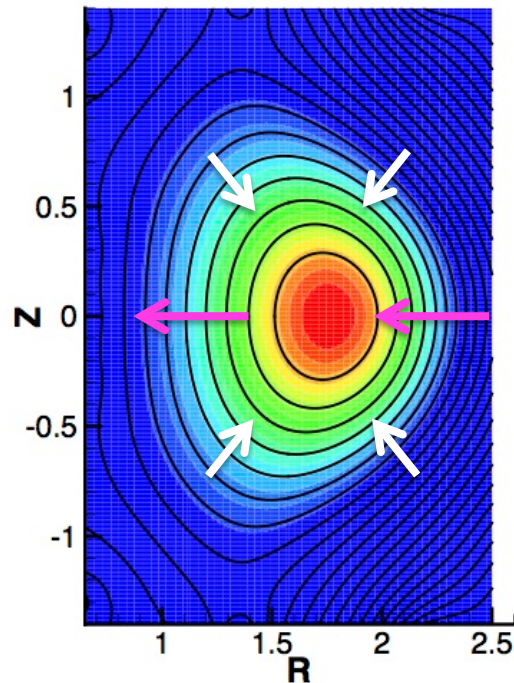
Background: Physical properties of confined plasma influence the selection of numerical methods.

- Separation of scales (medium sized tokamak)
 - Global Alfvén propagation time $\tau_A \sim 0.3 \mu\text{s}$
 - Particle collision times $\tau_e \sim 0.2 \text{ ms}$, $\tau_i \sim 15 \text{ ms}$
 - Global resistive diffusion time $\tau_r \sim 1\text{-}10 \text{ s}$ ($S = \tau_r / \tau_A$)
 - Sound gyroradius $\rho_s = (m_i k_B T_e / q_i^2 B^2)^{1/2} \sim 5 \text{ mm}$
 - Minor radius $a \sim 0.5 \text{ m}$
- Extreme anisotropy relative to $\mathbf{B}(\mathbf{x}, t)$
 - Effective thermal diffusivity ratio $\chi_{\parallel} / \chi_{\perp} \gg 10^6$
 - Also extreme anisotropy for viscous diffusivities
- Nonlinear conditions remain close to force-balance
 - No shock propagation
 - Distinct force-density contributions nearly cancel



Toroidal magnetic confinement has two primary sources of free energy for MHD.

- Toroidal geometry avoids end losses.
- \mathbf{B} must twist to prevent net outward drifts.
- Field-lines trace-out flux surfaces.



Cross section of plasma pressure contours (color) and magnetic flux (black lines). κ from \mathbf{B}_{pol} (white) and from B_ϕ (magenta).

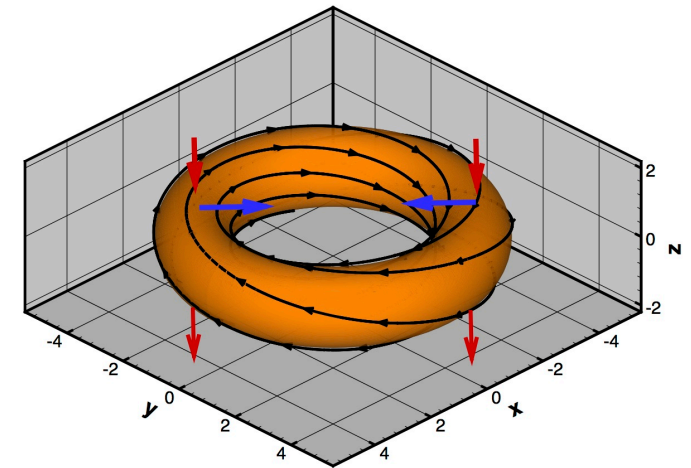


Illustration showing \mathbf{B} (black), ∇B vectors (blue) and resulting ion particle drifts (red).

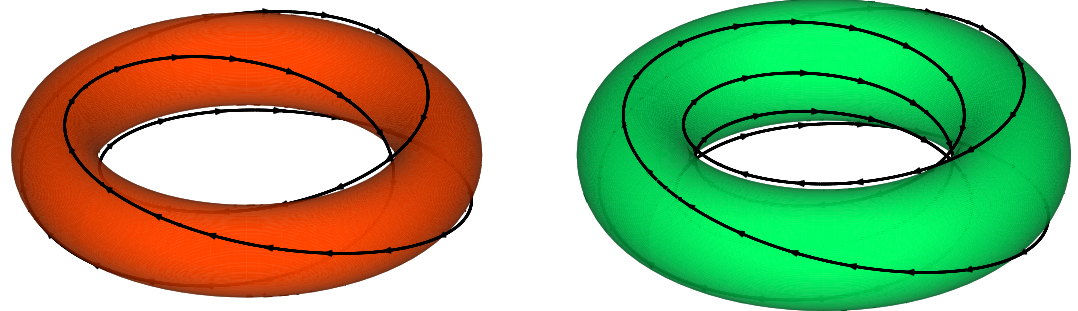
- #1) The alignment of curvature $\kappa = \hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}$ with ∇P leads to free energy; the perturbed ideal-MHD energy contribution is

$$- \int_{R_{pl}} \left(\vec{\xi}_\perp \cdot \nabla P \right) \left(\vec{\xi}_\perp^* \cdot \kappa \right) dVol \sim \mathbf{F} \cdot \Delta \mathbf{s}$$

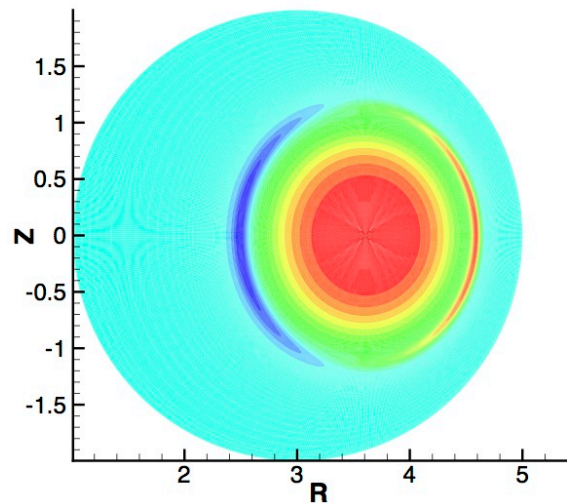


There are two primary sources of free energy (cont).

- In axisymmetric systems, twist is provided by charge current running through the plasma.
- Except in FRCs, current density is largely parallel to \mathbf{B} .



Two surfaces of the same configuration showing 3/2 (left) and 5/2 (right) twist, i.e. safety factor (q).



Contour plot of λ for the same configuration shows significant spatial variation.

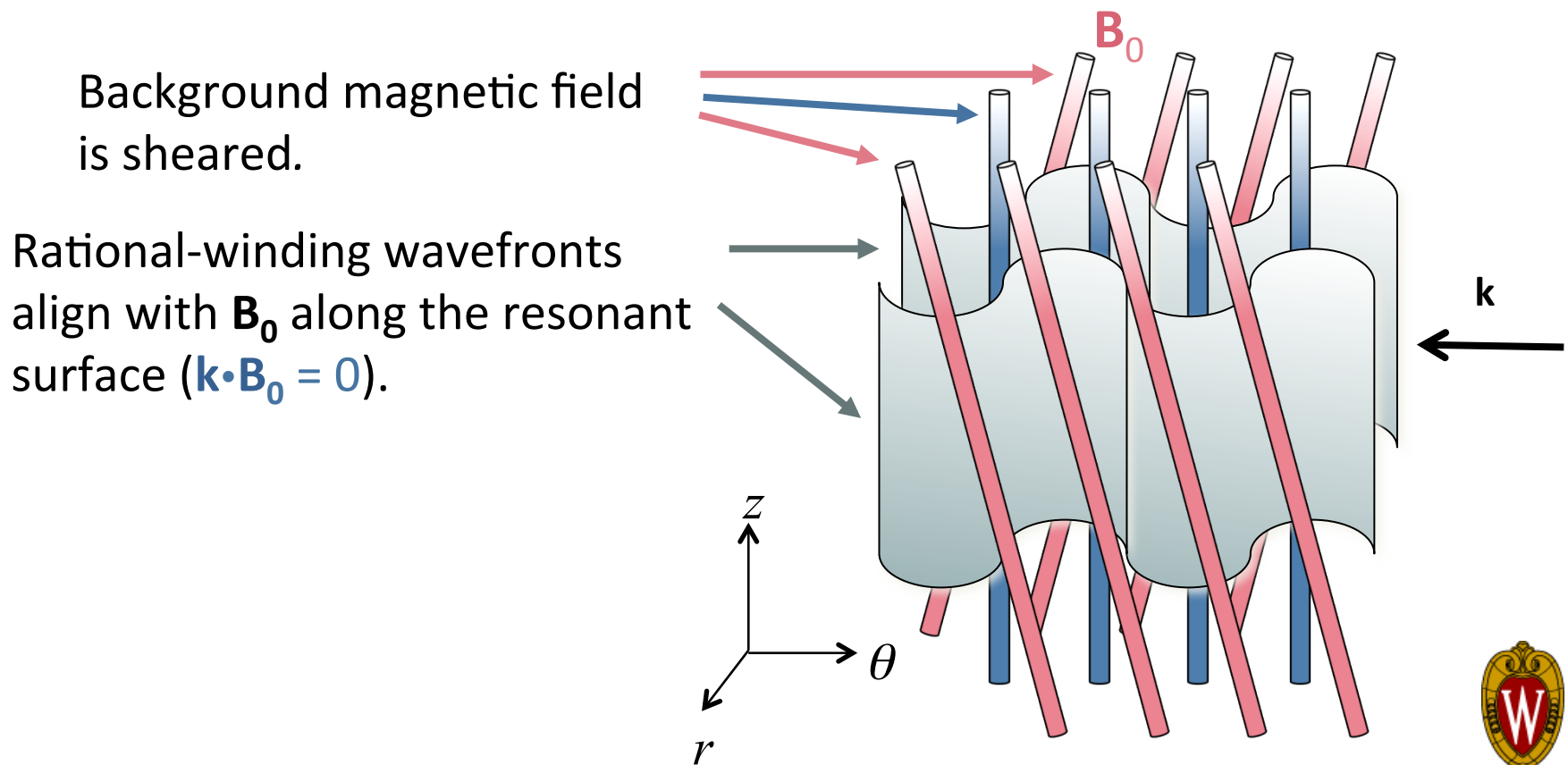
- #2) Spatial variation of the “parallel” current density, $\lambda = J_{\parallel}/B$, also contributes free energy:

$$- \int_{R_{pl}} \lambda \left(\vec{\xi}_{\perp}^* \times \mathbf{B} \right) \cdot \nabla \times \left(\vec{\xi}_{\perp} \times \mathbf{B} \right) dVol \sim IV \Delta t$$



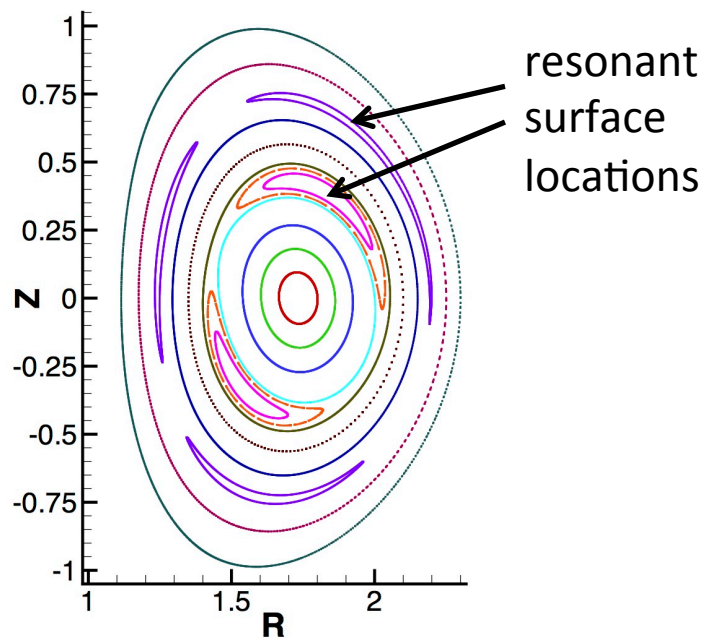
Resonance for helical shear-Alfvén waves occurs along toroidal surfaces.

- Restoring force density from bending is weak where wavefronts align with \mathbf{B}_0 , hence susceptibility to instability.
- Resonant instability in sheared field leads to spatial localization.

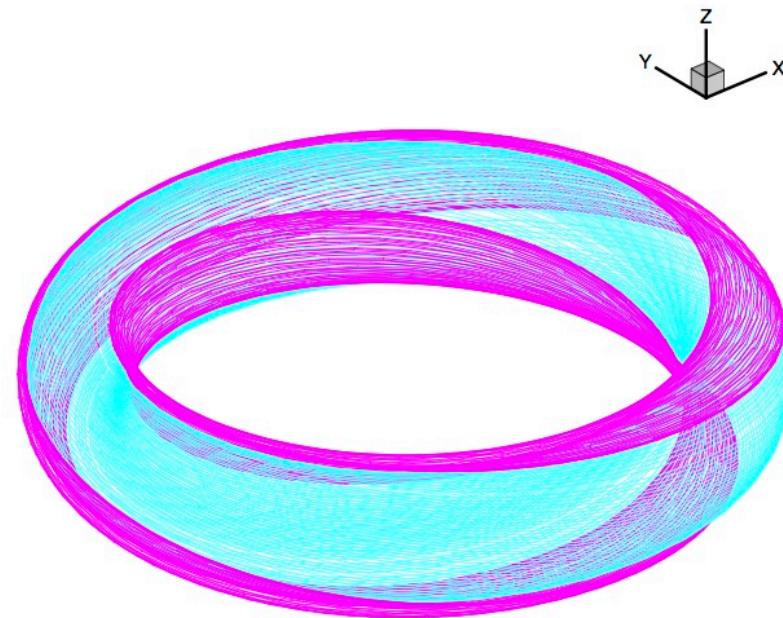


Pertinent examples: Non-ideal effects lead to changes in magnetic topology.

- Resistive or other non-ideal effects allow instabilities when free energy is insufficient for ideal-MHD instability.
- Magnetic reconnection (from $\nabla\lambda$ energy) at resonances leads to helical islands.
- Island overlap produces regions of stochastic magnetic field.



Cross-sections of islands are embedded among toroidal flux surfaces.

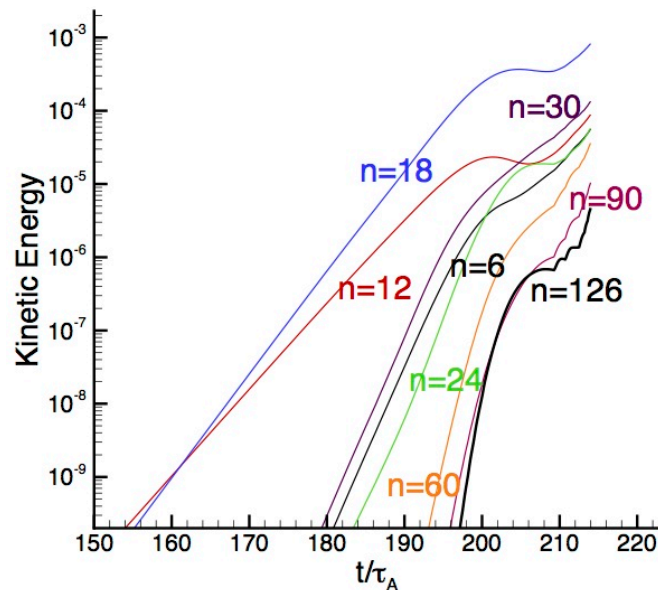


Non-overlapping islands are distinct regions but enhance energy transport.

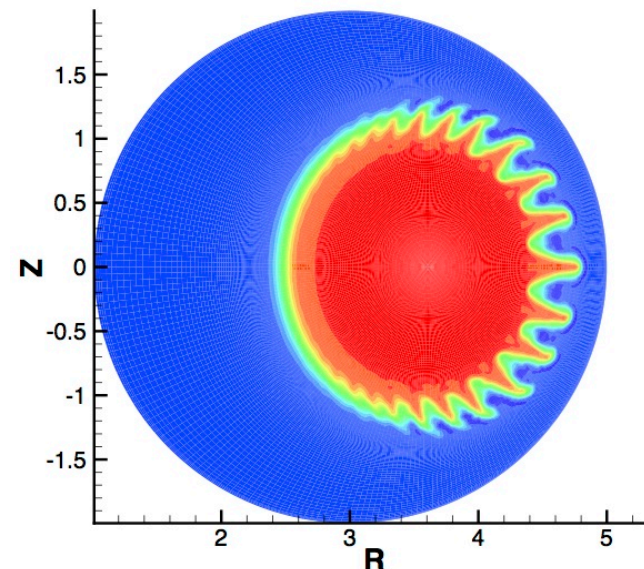


Interchange can localize on the outboard side of a torus, leading to ballooning instability.

- Like other interchange activity, dynamics are largely perpendicular to \mathbf{B} .
- Instability tends to arise over a broad range of toroidal wavenumbers; two-fluid and kinetic effects can be important.
- Ballooning can cause disruption or edge-localized modes (ELMs).



A computation run with limited periodicity has 2 linearly unstable wave-numbers.



Nonlinear evolution produces helical fingers of density and energy.



Models: We distinguish primitive-field and potential-field systems of equations.

- Primitive-field models describe the evolution of low-order moments of particle distributions and low-frequency electromagnetics.*

$$\frac{\partial n}{\partial t} + \nabla \cdot (n \mathbf{V}) = 0$$

particle continuity

$$mn \left(\frac{\partial}{\partial t} + \mathbf{V} \cdot \nabla \right) \mathbf{V} = \mathbf{J} \times \mathbf{B} - \nabla p - \nabla \cdot \underline{\Pi}$$

momentum density

$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E}$$

Faraday's law

$$\mu_0 \mathbf{J} = \nabla \times \mathbf{B}$$

Ampere's law

$$\nabla \cdot \mathbf{B} = 0$$

divergence constraint

- The model also needs closure information and a generalized Ohm's law for \mathbf{E} .

*See [Kimura and Morrison, PoP **21**, 082101] for energy considerations.



The remaining relations select the level of physics fidelity.

- The generalized Ohm's law extracts a low-frequency relation for \mathbf{E} from electron momentum-density evolution.

$$\mathbf{E} = \underbrace{-\mathbf{V} \times \mathbf{B}}_{\text{ideal MHD}} + \underbrace{\eta \mathbf{J}}_{\text{resistive E}} + \underbrace{\frac{1}{ne}(\mathbf{J} \times \mathbf{B} - \nabla p_e)}_{\text{Hall and } e^- \text{ pressure}} + \underbrace{\frac{1}{\epsilon_0 \omega_{pe}^2} \left[\frac{\partial}{\partial t} \mathbf{J} + \nabla \cdot (\mathbf{J} \mathbf{V} + \mathbf{V} \mathbf{J}) \right]}_{e^- \text{ inertia}}$$

- Stress may be a combination of effects.

$$\underline{\mathbf{W}} \equiv \nabla \mathbf{V} + \nabla \mathbf{V}^T - \frac{2}{3} \underline{\mathbf{I}} \nabla \cdot \mathbf{V}$$

$$\underline{\Pi}_{\text{gv}} = \frac{m_i p_i}{4eB} \left[\hat{\mathbf{b}} \times \underline{\mathbf{W}} \cdot (\underline{\mathbf{I}} + 3\hat{\mathbf{b}}\hat{\mathbf{b}}) - (\underline{\mathbf{I}} + 3\hat{\mathbf{b}}\hat{\mathbf{b}}) \cdot \underline{\mathbf{W}} \times \hat{\mathbf{b}} \right] \quad \text{gyroviscosity}$$

$$\underline{\Pi}_{\parallel} = \frac{p_i \tau_i}{2} (\hat{\mathbf{b}} \cdot \underline{\mathbf{W}} \cdot \hat{\mathbf{b}}) (\underline{\mathbf{I}} - 3\hat{\mathbf{b}}\hat{\mathbf{b}}) \quad \text{parallel}$$

$$\underline{\Pi}_{\perp} \sim -\frac{3p_i m_i^2}{10e^2 B^2 \tau_i} \underline{\mathbf{W}} \Rightarrow -nm_i v_{iso} \underline{\mathbf{W}} \text{ or } -nm_i v_{kin} \nabla \mathbf{V} \quad \text{perpendicular}$$



The closure relation for pressure(s) is selected for the dynamics of interest.

- At sufficiently low plasma- β ($= \mu_0 p / B^2$), pressure can be dropped.

$$p = 0$$

- If compressive waves are faster than all dynamics of interest, flow may be incompressible.

$$\nabla \cdot \mathbf{V} = 0$$

- An adiabatic relation describes fast perpendicular dynamics.

$$\frac{\partial}{\partial t} p + \mathbf{V} \cdot \nabla p = -\Gamma p \nabla \cdot \mathbf{V}$$

- Otherwise, an energy equation with heat-flux-density closure is used.

$$\frac{n_s}{\Gamma - 1} \left(\frac{\partial}{\partial t} T_s + \mathbf{V}_s \cdot \nabla T_s \right) = -n_s T_s \nabla \cdot \mathbf{V}_s - \nabla \cdot \mathbf{q}_s + Q_s \quad s = i, e$$

$$\mathbf{q}_s = -n_s (\chi_{\parallel} - \chi_{\perp}) \hat{\mathbf{b}} \hat{\mathbf{b}} \cdot \nabla T_s - n_s \chi_{\perp} \nabla T_s + \frac{5n_s T_s}{2q_s B} \hat{\mathbf{b}} \times \nabla T_s$$



anisotropic conduction

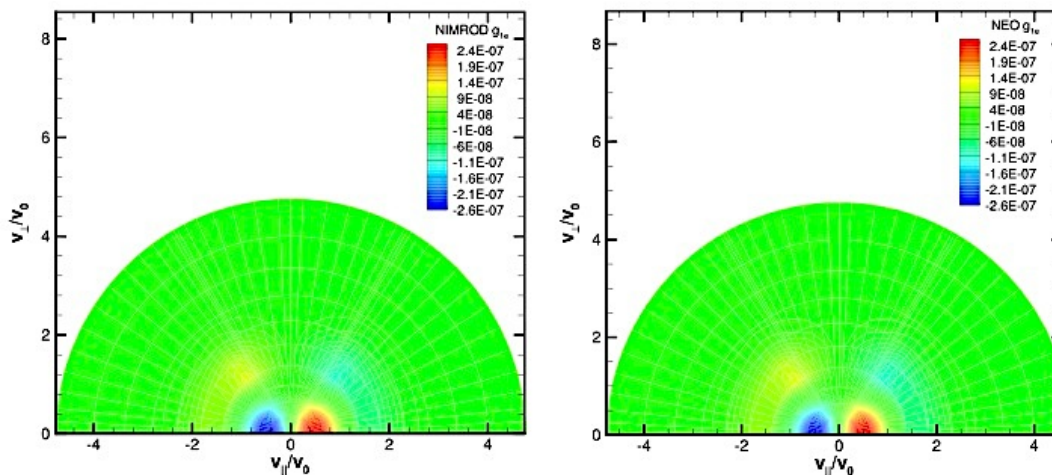


magnetization \mathbf{q}



Closure information can be obtained from kinetic computations.

- PIC computations have been used for a hot-ion component [Park, et al., PFB **4**, 2033; Kim, PoP **15**, 072507].
 - Low n_{hot} and $mn_{\text{hot}} \mathbf{V}_{\text{hot}}$ are assumed.
 - Coupling to \mathbf{V}_{COM} evolution is through hot-particle stress or current.
 - Applications include energetic-particle modes, sawteeth, and tearing.
- Eulerian δf drift-kinetics can be used to close majority-species equations.
 - Consistent kinetic model has been derived [Ramos, PoP **15**, 082106].
 - The kinetic computations have been solved and verified in the framework of an extended-MHD code [Held, et al., PoP **22** 032511].



Verification of the NIMROD implementation with NEO includes e^- distribution and bootstrap current [Held].



Potential formulations separate physical effects among the variables, themselves.

- The basic version is reduced MHD [Strauss, PoF **18**, 134], which orders tokamak fields by $\varepsilon = a/R$.
 - Twist of $O(1)$ implies $B_{pol} \sim \varepsilon B_\phi$ and $\beta \sim \varepsilon^2$.
 - Dynamics with $k_{||}/k_\perp \ll 1$ have $\mathbf{V}_\perp = \hat{\mathbf{b}}_0 \times \nabla \varphi$ at lowest order, where φ is the electrostatic potential.
 - The lowest-order perturbed field is $\mathbf{B}_1 = \nabla \psi \times \mathbf{B}_0$.
- A reduced resistive-MHD system is [Hazeltine and Meiss (1992)]:

$$\rho_0 \frac{\partial}{\partial t} \nabla_f^2 \varphi = - \frac{B_0^2}{\mu_0} \nabla_{||} \nabla_f^2 \psi + 2 \hat{\mathbf{b}}_0 \times \boldsymbol{\kappa} \cdot \nabla_f p \quad \text{parallel vorticity evolution}$$

$$\frac{\partial}{\partial t} \psi = - \frac{1}{B_0} \nabla_{||} (B_0 \varphi) + \frac{\eta}{\mu_0} \nabla_f^2 \psi \quad \text{parallel Ohm's law}$$

$$\frac{\partial}{\partial t} p = \hat{\mathbf{b}}_0 \times \nabla p_0 \cdot \nabla_f \varphi \quad \text{pressure advection}$$

- For $\varepsilon \ll 1$, $||$ is the toroidal direction and ∇_f is the poloidal gradient.



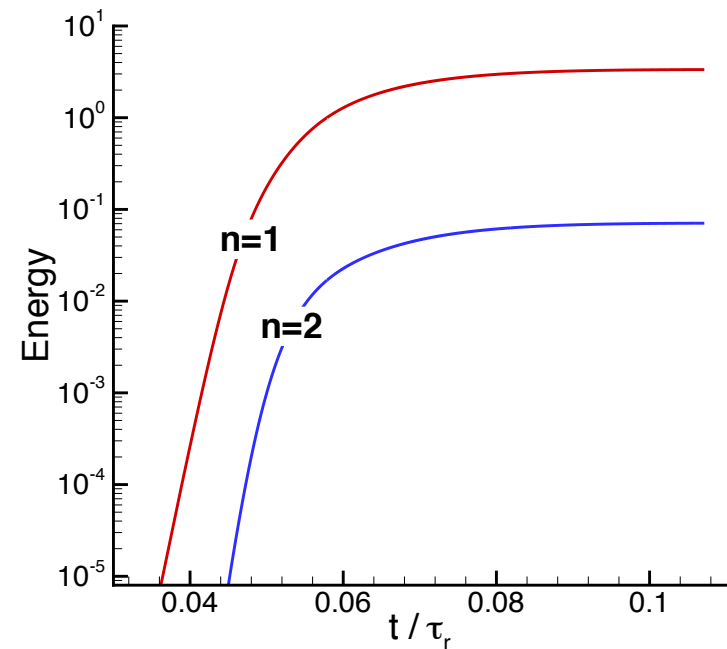
The choice of model has ramifications for numerical computation.

- Potential-based systems [e.g. Breslau, Ferraro, Jardin, PoP **16**, 092503]
 - Numerical operations on scalars are relatively straightforward.
 - Potentials can avoid numerical coupling of physical effects.
 - Higher-order differentiation or use of auxiliary variables is required.
 - Non-reduced equations are more complicated than the vector moment relations.
- Primitive-field systems [e.g. Sovinec, et al., JCP **195**, 355]
 - Vector calculus is not trivial, numerically.
 - Separating distinct physical effects relies on numerics.
 - Equations are directly from moment relations.
 - Representations can use lower-order continuity.



Numerical methods (part 1): Some form of implicit computation is needed.

- Nonlinear evolution occurs over long time-scales.
 - Virulent ELMs occur over 10s of τ_A .
 - Resistive island evolution and relaxation occur over tenths of τ_r .
- Wave-CFL condition restricts explicit computation.
 - Resolving resonances requires $\Delta x < \rho_s$.
 - $\Delta x < R/1000$ implies 1000s of steps for each τ_A of an explicit simulation.
- A number of implicit methods have been applied to MHD for magnetic confinement.
 - Many fall into the class of methods now labeled “IMEX” in applied math.



Development of magnetic island shown earlier occurs over $\sim 10^5 \tau_A$.



A simple example illustrates practical alternatives.

- Take the cylindrical-geometry approximation with uniform $\mathbf{B}_0 = B_0 \hat{\mathbf{z}}$ and vanishingly small β .
 - The ∇_f^2 operator can be inverted.
 - In the simplified, reduced system with $\eta = 0$, small perturbations of the form $f(r)\exp[im\theta + ikz]$ evolve according to

$$\frac{\partial}{\partial t} \varphi_{m,k} = -v_A^2 \nabla_{\parallel} \psi_{m,k} \rightarrow -ikv_A^2 \psi_{m,k}$$

$$v_A^2 = B_0^2 / \mu_0 \rho_0$$

$$\frac{\partial}{\partial t} \psi_{m,k} = -\nabla_{\parallel} \varphi_{m,k} \rightarrow -ik\varphi_{m,k}$$



Explicit methods are limited by the *CFL* condition.

- Apply an explicit leapfrog time-advance ($t^n = n\Delta t$) to the simplified shear-wave system:

$$\begin{aligned}\varphi_{m,k}^{n+1} - \varphi_{m,k}^n &= -ik\Delta t v_A^2 \psi_{m,k}^n \\ \psi_{m,k}^{n+1} - \psi_{m,k}^n &= -ik\Delta t \varphi_{m,k}^{n+1}\end{aligned}$$

- Taking ξ as the eigenvalue of the time-step operation,

$$\begin{aligned}(\xi - 1)\varphi_{m,k} &= -ik\Delta t v_A^2 \psi_{m,k} \\ (\xi - 1)\psi_{m,k} &= -ik\Delta t \xi \varphi_{m,k}\end{aligned}$$

- Solutions of the characteristic equation, $(\xi - 1)^2 + k^2 v_A^2 \Delta t^2 \xi = 0$, have

$$\begin{aligned}|\xi| > 1, \Delta t > \frac{2}{kv_A} &\rightarrow \text{unphysical growth (linear numerical instability)} \\ |\xi| = 1, \Delta t \leq \frac{2}{kv_A} &\rightarrow \text{stable oscillation}\end{aligned}$$



Implicit methods allow numerical stability at large time-step values.

- A flexible method evaluates the drive terms at f into the step ($0 \leq f \leq 1$):

$$\varphi_{m,k}^{n+1} - \varphi_{m,k}^n = -ik\Delta t v_A^2 \left[f\psi_{m,k}^{n+1} + (1-f)\psi_{m,k}^n \right]$$

$$\psi_{m,k}^{n+1} - \psi_{m,k}^n = -ik\Delta t \left[f\varphi_{m,k}^{n+1} + (1-f)\varphi_{m,k}^n \right]$$

- Taylor-expanding the analytical solution about $t^n + \frac{1}{2}\Delta t$ and inserting into the approximation, to $O(\Delta t^2)$,

$$\left(\frac{\partial}{\partial t} + \frac{\Delta t^2}{24} \frac{\partial^3}{\partial t^3} \right) \begin{pmatrix} \varphi_{m,k} \\ \psi_{m,k} \end{pmatrix} = -ik \left[1 + \Delta t \left(f - \frac{1}{2} \right) \frac{\partial}{\partial t} + \frac{\Delta t^2}{8} \frac{\partial^2}{\partial t^2} \right] \begin{pmatrix} v_A^2 \psi_{m,k} \\ \varphi_{m,k} \end{pmatrix}$$

- Keeping the lowest-order terms, the differential approximation [Shokin (1983)] can be expressed as

$$\frac{\partial}{\partial t} \begin{pmatrix} \varphi_{m,k} \\ \psi_{m,k} \end{pmatrix} = -ik \begin{pmatrix} v_A^2 \psi_{m,k} \\ \varphi_{m,k} \end{pmatrix} - \Delta t \left(f - \frac{1}{2} \right) k^2 v_A^2 \begin{pmatrix} \varphi_{m,k} \\ \psi_{m,k} \end{pmatrix}$$

★ k^2 represents $-\nabla_{\parallel}^2$, so the last term adds numerical damping for $f > \frac{1}{2}$.



Researchers have used alternatives to full implicit methods.

- Implicit computations solve algebraic systems at each step.
 - Coefficients of the spatial expansion are coupled at the new time.
 - Full implicit methods solve nonlinear algebraic systems.
 - System size, condition number, sparsity, and linear vs. nonlinear affect computational cost per step.
- The “quasi-implicit” method uses the large- R/a ordering and treats only poloidal compression implicitly [Park and Monticello, NF **30**, 285].
 - The potential formulation separates poloidal compression to keep algebraic systems small and linear (original M3D).
 - Shear waves (explicit) diminish computational gains at moderate R/a .
- “Semi-implicit” methods add numerical dispersion to stabilize the advance [Schnack, et al., JCP **70**, 330] (DEBS, original XTOR, NIMROD).



The semi-implicit leapfrog is compatible with primitive-field representations.

- Add a positive spatial differential operator $-\Delta t^2 \underline{\mathbf{L}}$ to staggered leapfrog for linear ideal-MHD--dropping continuity and pressure for clarity:

$$\left(\rho_0 - \Delta t^2 \underline{\mathbf{L}}\right) \left(\mathbf{V}^{n+1} - \mathbf{V}^n\right) = \Delta t \left(\mathbf{J}^{n+1/2} \times \mathbf{B}_0 + \mathbf{J}_0 \times \mathbf{B}^{n+1/2}\right)$$

$$\mathbf{B}^{n+1/2} - \mathbf{B}^{n-1/2} = \Delta t \nabla \times \left(\mathbf{V}^n \times \mathbf{B}_0\right)$$

- The differential approximation for the original initial conditions [Caramana, JCP **96**, 484] and with synchronization [Sovinec & King, JCP **229**, 5803] is

$$\left(\rho_0 - \Delta t^2 \underline{\mathbf{L}}\right) \frac{\partial}{\partial t} \mathbf{V} \Big|_{t^{n+1/2}} = \frac{1}{\mu_0} \nabla \times \left(\mathbf{B} + \frac{\Delta t}{2} \frac{\partial}{\partial t} \mathbf{B} \right) \Big|_{t^n} \times \mathbf{B}_0 + \mathbf{J}_0 \times \left(\mathbf{B} + \frac{\Delta t}{2} \frac{\partial}{\partial t} \mathbf{B} \right) \Big|_{t^n}$$

$$\frac{\partial}{\partial t} \mathbf{B} \Big|_{t^n} = \nabla \times \left[\left(\mathbf{V} - \frac{\Delta t}{2} \frac{\partial}{\partial t} \mathbf{V} \right) \Big|_{t^{n+1/2}} \times \mathbf{B}_0 \right]$$



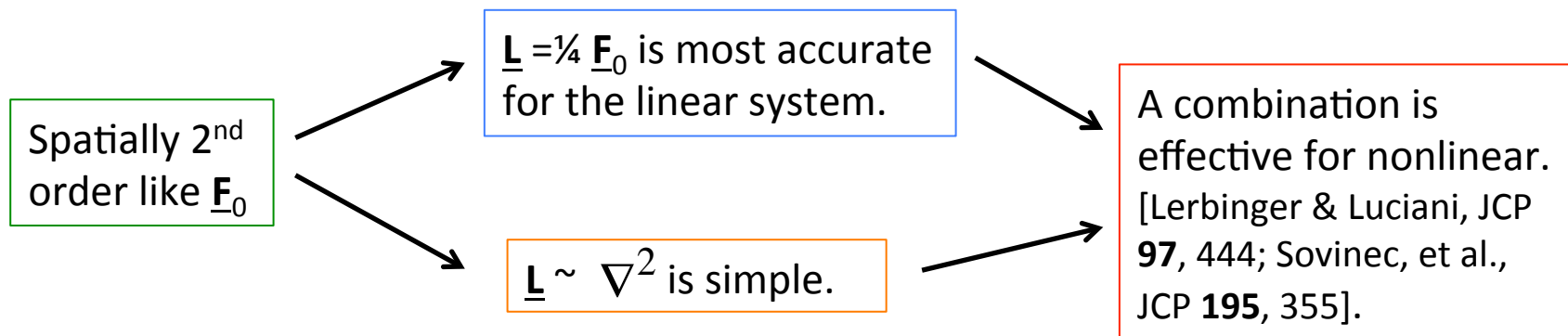
Manipulating the differential approximation shows the numerical properties of the semi-implicit method.

- The equations can be combined to produce the second-order

$$\left(\rho_0 + \frac{\Delta t^2}{4} \underline{\mathbf{F}}_0 - \Delta t^2 \underline{\mathbf{L}} \right) \frac{\partial^2}{\partial t^2} \mathbf{V} = \underline{\mathbf{F}}_0(\mathbf{V})$$

where $\underline{\mathbf{F}}_0$ is the linear ideal-MHD force operator.

- Modes of the linear ideal-MHD system satisfy $\underline{\mathbf{F}}_0(\xi) = -\rho_0 \omega^2 \xi$.
- In the absence of the semi-implicit operator, the system is ill-posed, i.e. numerically unstable if $\Delta t^2 > 4/\omega^2$ for the largest ω^2 .
- $\underline{\mathbf{L}}$ can be selected for accuracy and computational practicality:

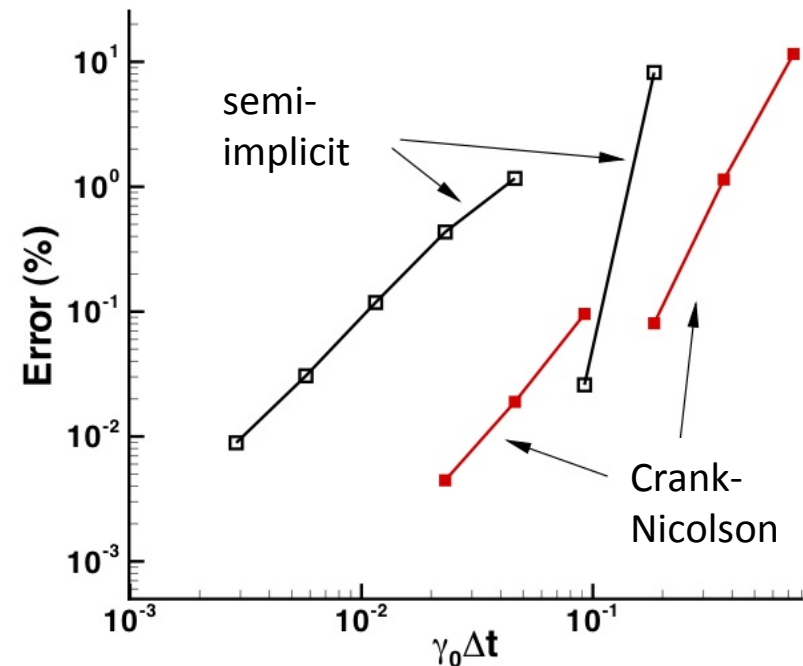


Implicitly balanced methods avoid truncation errors from separating (“splitting”) physics.

- Quasi- and semi-implicit methods use various levels of splitting.
- Balanced methods determine all fields at the new time-level, simultaneously.
 - This improves multi-scale convergence [Knoll, et al. JCP **185**, 583].
 - A simple analysis for $du/dt = \gamma_0 u$ with Crank-Nicolson (implicit $f = 1/2$) is

$$\gamma_{CN} = \frac{\ln(\xi_{CN})}{\Delta t} = \frac{1}{\Delta t} \ln\left(\frac{2 + \gamma_0 \Delta t}{2 - \gamma_0 \Delta t}\right)$$

- Modern Krylov-space algebraic solvers + Newton’s method facilitate balanced nonlinear MHD computation [Chacón, PoP **15**, 056102].
- Avoiding numerical dissipation is important for simulating high temperatures.

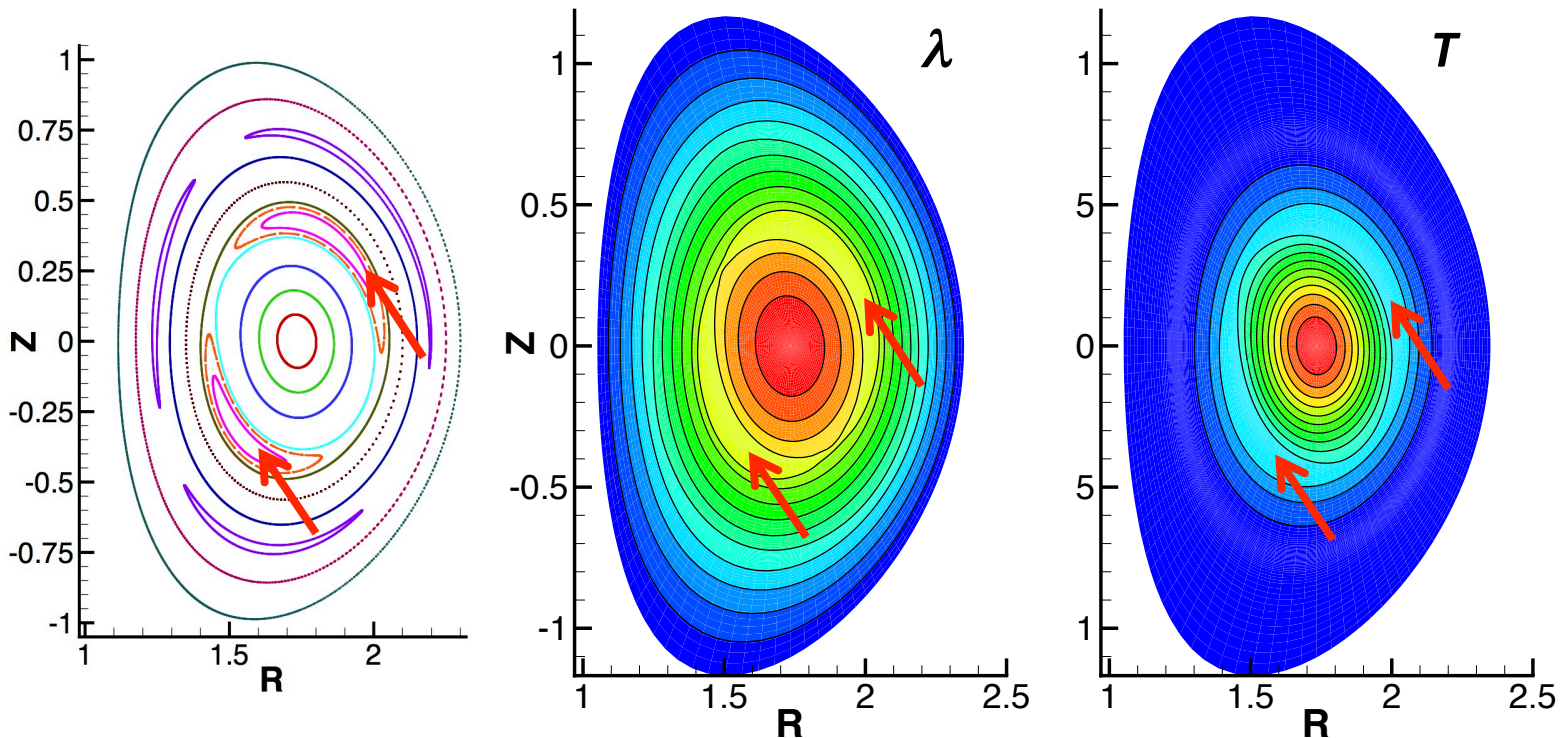


Tearing-mode comparison of C-N and SI [JCP **229**, 5803] for two-fluid.



Numerical methods (part 2): There are several considerations for spatial representations.

- Accurate representation of $\hat{\mathbf{b}} \cdot \nabla$ is important for force equilibration.
- Anisotropic transport also depends on the evolving $\hat{\mathbf{b}} \cdot \nabla$.

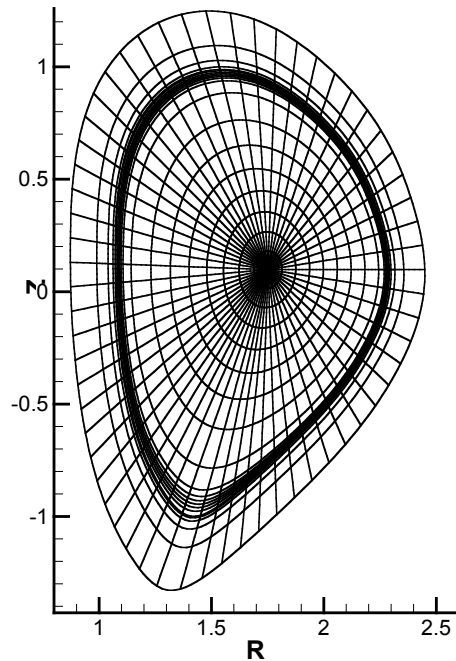


The island-evolution example saturates nonlinearly by making λ uniform along \mathbf{B} , and strongly anisotropic $\underline{\kappa}$ equilibrates T over the island.

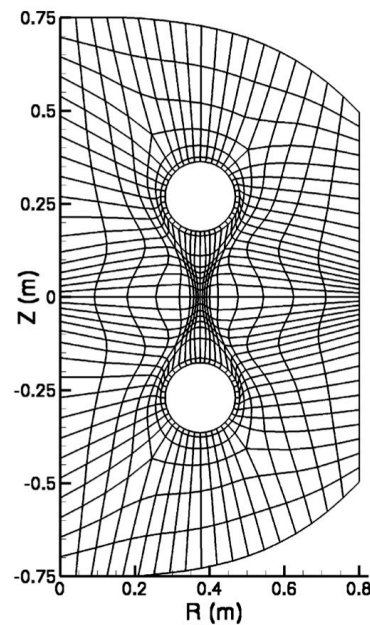


Geometry is important for simulating macroscopic dynamics in experiments.

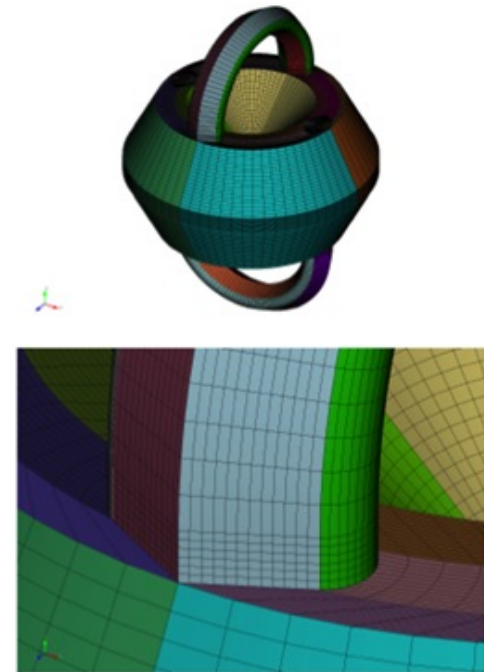
- Finite Fourier series for periodic coordinates is effective but geometrically limiting.
- Meshing two or all three coordinates enhances flexibility.



Packed mesh of curved 2D elements for modeling ELMS.



2D mesh for modeling MRX [Murphy, PoP **15**, 042313].

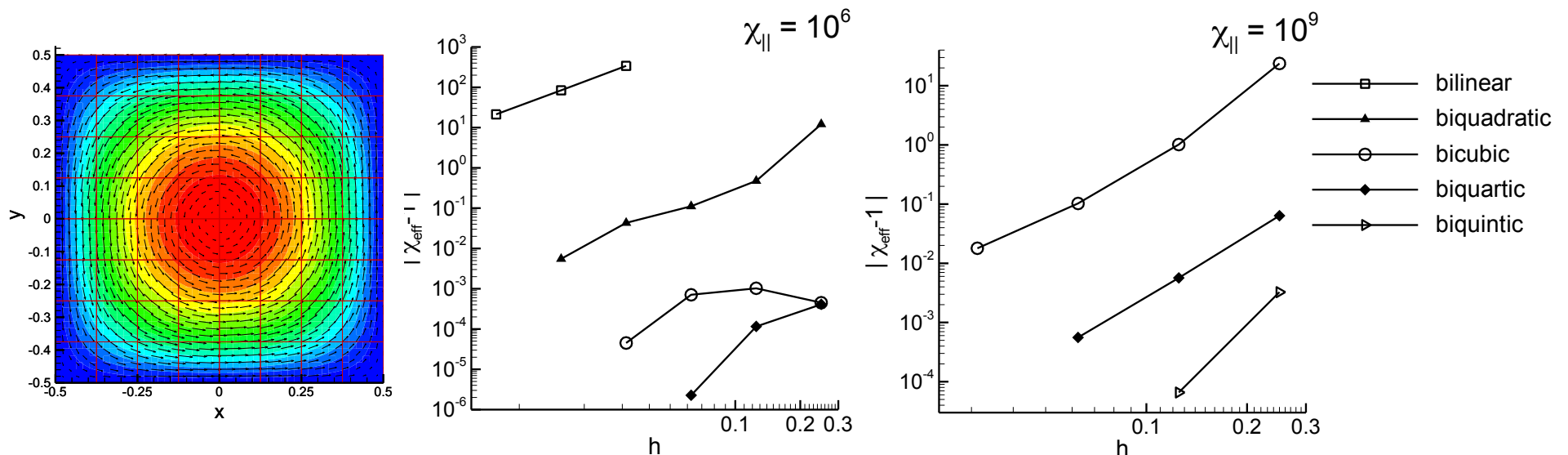


CAD-based 3D mesh for HIT-SI. [PSI-Center and CUBIT]



High-order elements and spectral representations are effective for anisotropic transport.

- A thermal-conduction test with analytical solution quantifies transport error.
- Results with high-order elements converge on extreme anisotropy without mesh alignment [Sovinec, et al., JCP **195**, 355].



The test case has magnetic flux and $T \sim \cos(\pi x)\cos(\pi y)$.

Numerical error in perpendicular transport is quantified by the computed peak temperature as mesh size and polynomial bases are varied.



Element-based function spaces need to be suited for the system of equations.

- Dependent variables are expanded in polynomials within each element.
- Projections generate algebraic equations from the differential system.
- A 1D thermal-conduction example illustrates the process:

$$\text{Solve} \quad -\frac{d}{dx}\kappa(x)\frac{dT}{dx} = Q(x) \quad \text{in} \quad 0 \leq x \leq L$$

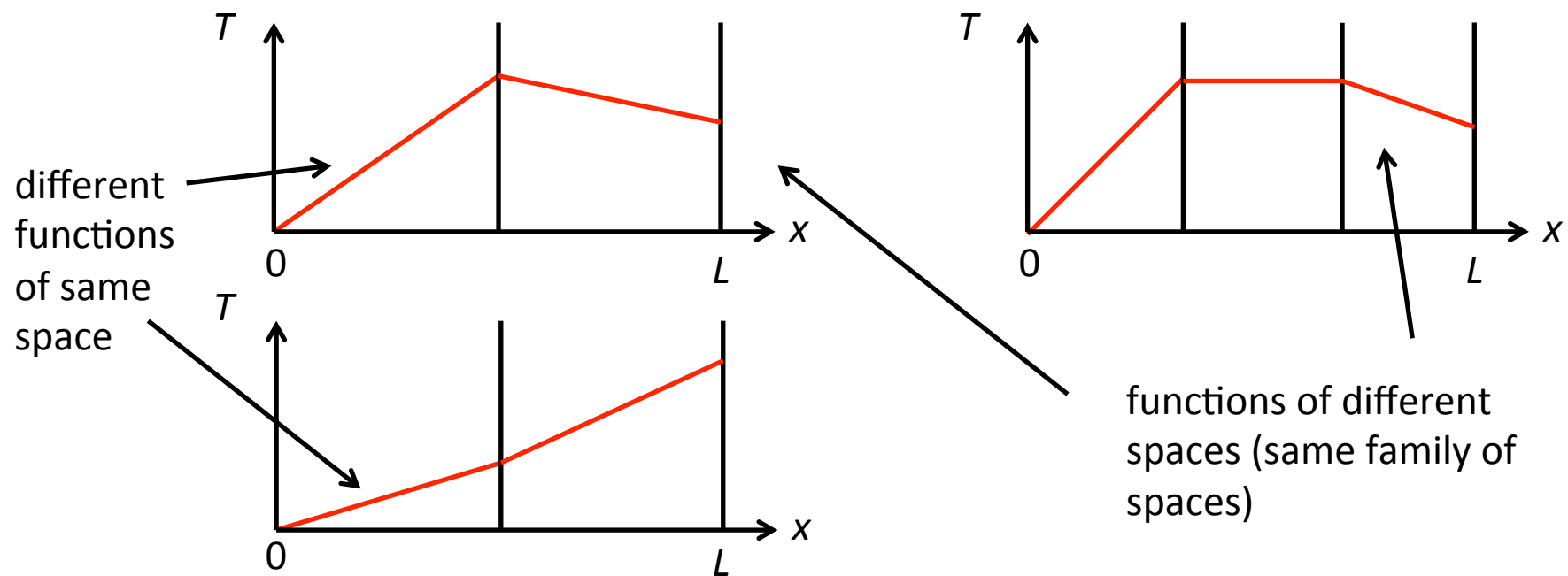
$$\text{subject to} \quad T(0) = T_0, \quad -\kappa(L)\frac{\partial T}{\partial x} = q_L$$

assuming $\kappa(x) > 0$.

- The formal solution $T(x) = T_0 + \int_0^x \frac{1}{\kappa(x'')} \left[\int_{x''}^L Q(x') dx' - q_L \right] dx''$ has continuous $T(x)$, but dT/dx is discontinuous at jumps in κ or point sources in Q .
- The suitable function space has C^0 continuity.



Finite-dimensional function spaces are defined by the choice of mesh and polynomials within each element.



- The above sketches of linear elements illustrate finite-dimensional function spaces of C^0 continuity.
- $T(x)$ is a nodal expansion:
$$T^h(x) = \sum_i T_i \alpha_i(x)$$



The weak form of an equation is used to select the best function from a given space.

- For a given mesh and basis, find $T^h(x) \in S^h$ such that

$$\int_0^L \left[\kappa \frac{d\Theta}{dx} \frac{dT^h}{dx} - \Theta Q \right] dx - q_L \Theta(L) = 0$$

for all $\Theta(x) \in S^h$.

- The integrals generate an algebraic system for the coefficients of T^h , symbolically expressed as

$$\underline{\underline{MT}} = \underline{R}$$

- The first term in the integral is a mathematical energy that responds to all possible wiggles in the function space.

*See Strang and Fix, *An Analysis of the Finite Element Method* (1988).



There are spatial-representation challenges for MHD applications.

- Dissipative terms are second-order, so continuity requirements are similar to the conduction example.

➤ However, dissipation is weak in high temperature plasma.

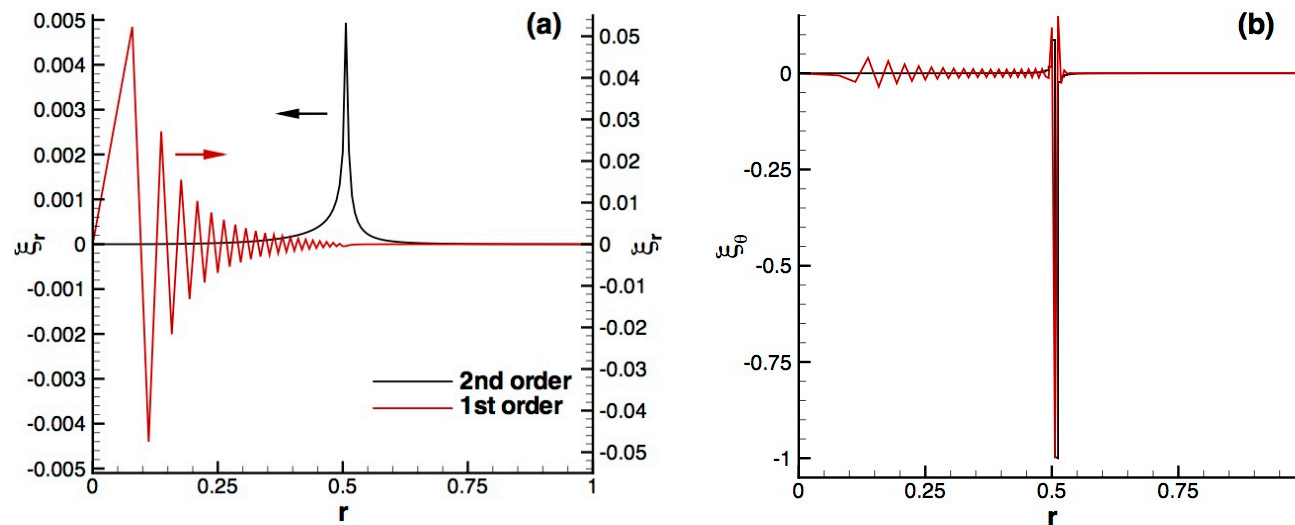
$$R_{cell} = \frac{\Delta x |\mathbf{V}|}{\nu} \begin{matrix} > \\ < \end{matrix} 1? \quad S_{cell} = \frac{\Delta x v_A \mu_0}{\eta} \begin{matrix} > \\ < \end{matrix} 1?$$

- The ideal part of the primitive-field time-dependent equations have first-order spatial derivatives.
 - Galerkin projection does not respond to all wiggles.
- The ∇_{\parallel} operator is singular for a helical distortion along its surface of resonance.
 - Free energy from bad curvature can excite mesh-scale oscillations.
- Satisfying the divergence constraint is not trivial for expansions of \mathbf{B} .
- Unlike the simple example, identifying appropriate function-spaces for MHD and extended-MHD is not easy.



Consequences of a poor choice can be significant.

- Spectral pollution and numerical destabilization of physically stable interchange are well known concerns.
 - 2nd-order ideal-MHD eigenvalue problems: [Gruber and Rappaz (1985); Degtyarev and Medvedev, CPC **43**, 29]
 - 1st-order t -dependent: [Lütjens and Luciani, CPC **95**, 47; Sovinec, JCP **319**, 61]



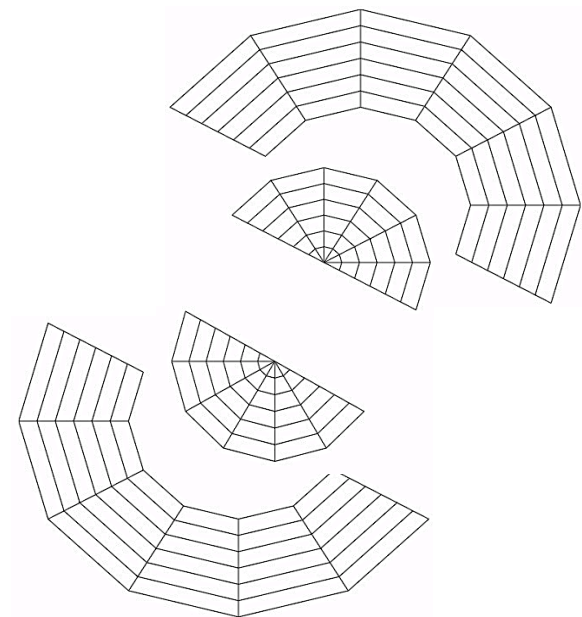
Components of a) radial and b) azimuthal displacement from 2nd-order (black) and unstabilized 1st-order (red) computations of marginally stable interchange.

- Physically representative behavior requires appropriate numerical responses to singular bending and compression at the limit of resolution.



Parallel computing: 3D domain decomposition is needed for large nonlinear problems.

- Domain decomposition is straightforward for element-based methods of generating algebraic systems.
- Codes using 3D elements, e.g. M3D-C1, decompose geometrically over all dimensions.
- Codes using finite Fourier series for one or two dimensions, e.g. NIMROD, decompose those dimensions by Fourier component.
- ✧ Solving the algebraic systems from implicit advances dominates parallel performance.
 - CG and GMRES operations scale well, but they need preconditioning.
 - Multigrid has optimal scaling but does not work well on all matrices.

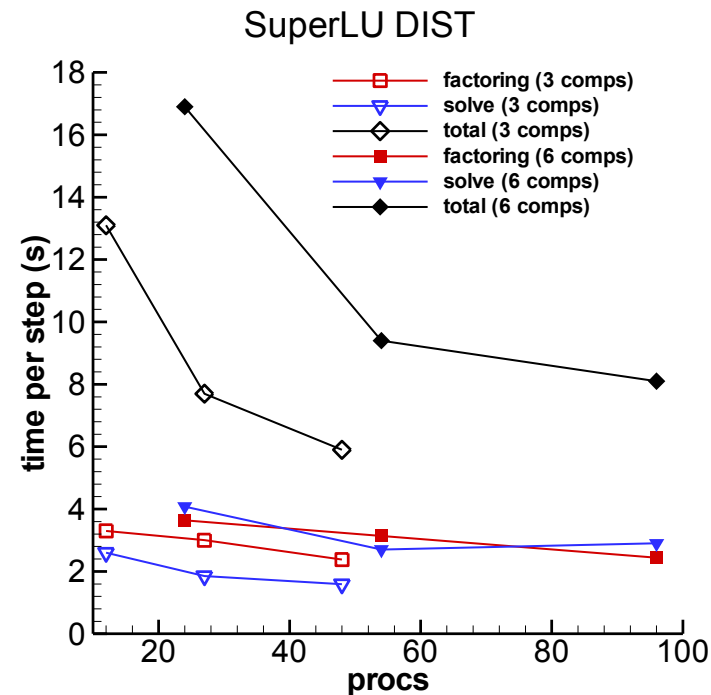
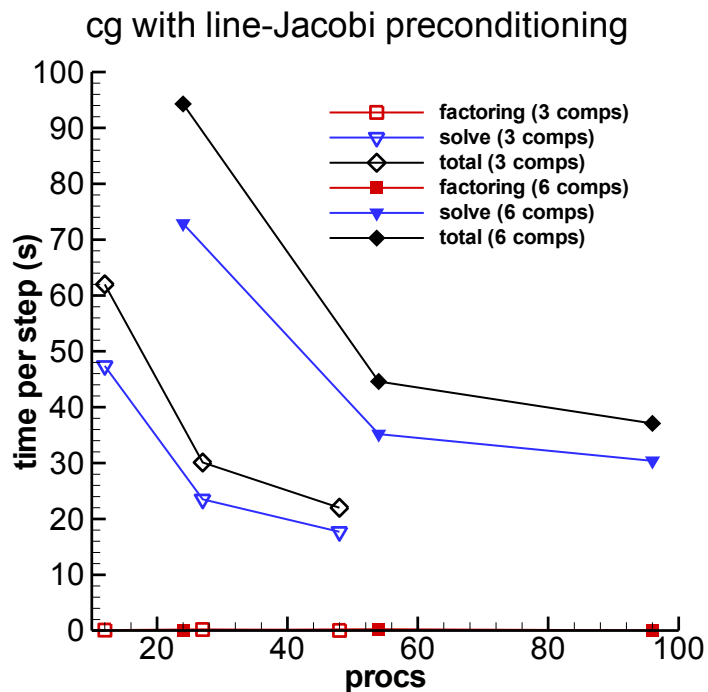


Mesh of 2D elements decomposed into blocks for parallel processing.



Distributed-memory parallelism with MPI communication has been the standard.

- The parallel performance of the preconditioning operations influences scaling and overall speed.



Strong-scaling tests on a small problem compare a less-effective preconditioner (left) with parallel sparse solves of diagonal blocks (right) [SuperLU: Li & Demmel ACM TMS **29**, 110].

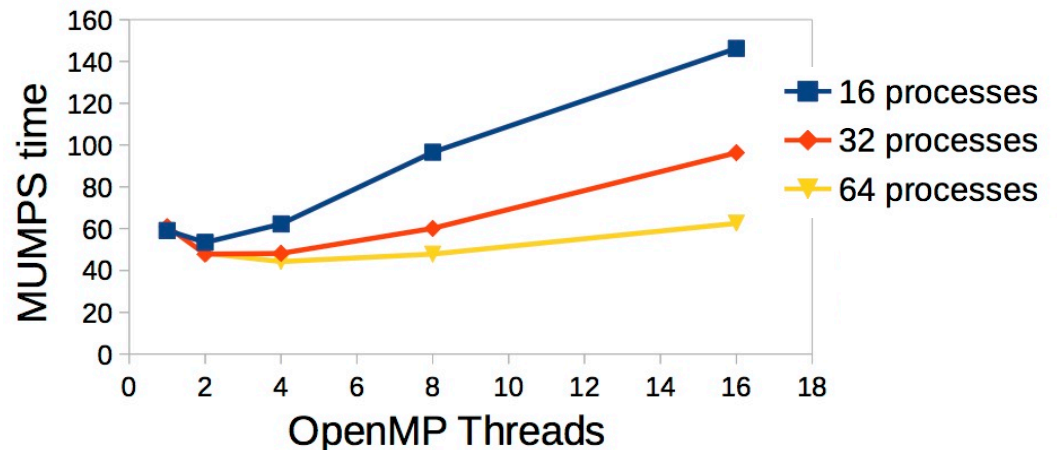
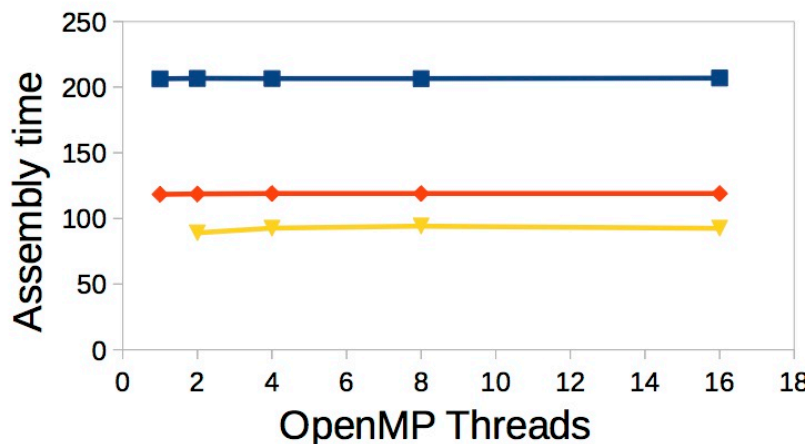


Fusion MHD computing is gradually evolving from MPI to hybrid parallelization.

- Efficient use of on-node memory benefits from thread-based parallelism.
- Obtaining performance improvement requires directives, e.g. OpenMP.
- Jacob King, Tech-X, implemented and tested OpenMP in NIMROD.

Mira BlueGeneQ at ORNL has symmetric multi-processing (SMP) nodes.

- Each node has 16 cores.
- Hyper-threading *allows* over-subscription by up to 4.



King's single-node tests with NIMROD + MUMPS 5.0.0 show:

- Exchanging MPI calls with OpenMP threads has little impact for FE “assembly.”
- MUMPS 5.0.0 threading is better than MPI alone (1 vs. 2 threads).



Open Challenges and Outlook

- “Multi-physics” challenge: applications, such as tokamak disruption, requires more general physics modeling.
 - Electromagnetics interact with 3D external conductors.
 - Plasma-surface interaction affects plasma properties.
 - Neutral dynamics and radiation are important.
 - Runaway e^- form a significant new kinetic species.
- “Multi-scale” challenge: temporal and spatial scale separation remain primary applied-math considerations.
 - Ranges of scales in experiments increase with plasma performance.
 - Drift physics (even 2-fluid) introduces oscillations -> implicit $\Delta t > \omega^{-1}$??
 - SOLVERS, SOLVERS, SOLVERS
- Increasing hardware complexity is a challenge for implicit computation.
 - Frequent data movement is needed for implicit computation.
 - Solvers??



Addressing the challenges will need plasma theory, applied mathematics, and computer science.

- Code-coupling is often assumed to be the fastest approach to multi-physics simulation.
 - Implicit stepping may need outer iteration.
 - Coupling computations that use different representations needs more study.
- Hardware accelerators, e.g. GPUs, compounds data movement problems.
 - Computer-science development of algebraic solvers is needed.
 - Use for spatially local computations including v-space computation?
 - Revisit less-implicit methods?
- ❖ Cross-disciplinary teaming will continue to be the best approach to meeting the challenges of fusion MHD simulation.



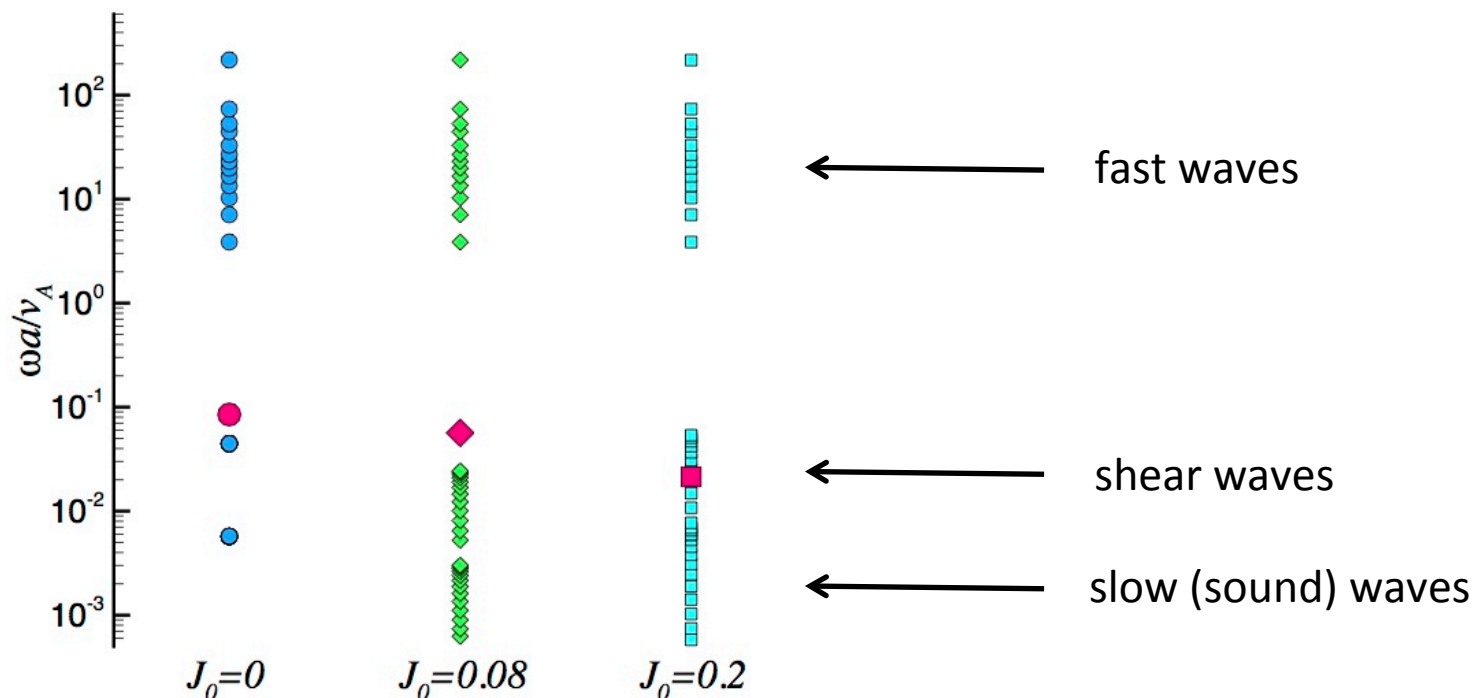
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The ideal-MHD spectrum of a periodic cylinder illustrates different types of modes.

- There is a gap between the uniform-density “plasma” column $0 \leq r \leq 1$ and a conducting wall at $r = 1.5$.
- Spectra for $m = 1$, $k = -0.0445$ with varying levels of parabolic axial current density are evaluated for $\mu_0 J_{0z} = J_0(1 - r^2)$ and $\beta(0) = 2\%$.



Magenta is an external mode that is unstable at $J_0 = 0.2$, where $\text{Im}(\omega)$ is shown.

